

# Black Probes of Schrödinger Spacetimes

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## Abstract

We consider black probes of Anti-de Sitter and Schrödinger spacetimes embedded in string theory and M-theory and construct perturbatively new black hole geometries. We begin by reviewing black string configurations in Anti-de Sitter dual to finite temperature Wilson loops in the deconfined phase of the gauge theory and generalise the construction to the confined phase. We then consider black strings in thermal Schrödinger, obtained via a null Melvin twist of the extremal D3-brane, and construct three distinct types of black string configurations with spacelike as well as lightlike separated boundary endpoints. One of these configurations interpolates between the Wilson loop operators, with bulk duals defined in Anti-de Sitter and another class of Wilson loop operators, with bulk duals defined in Schrödinger. The case of black membranes with boundary endpoints on the M5-brane dual to Wilson surfaces in the gauge theory is analysed in detail. Four types of black membranes, ending on the null Melvin twist of the extremal M5-brane exhibiting the Schrödinger symmetry group, are then constructed. We highlight the differences between Anti-de Sitter and Schrödinger backgrounds and make some comments on the properties of the corresponding dual gauge theories.

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## 1 Introduction

The in-depth study of Schrödinger spacetimes and its physical applications began with the recent work of [1, 2, 3, 4, 5], motivated by possible holographic applications to the description of condensed matter systems with non-relativistic symmetries.<sup>1</sup> The embedding of these spacetimes into type IIB string theory and their relation to the AdS/CFT correspondence via the null Melvin twist of known supergravity branes [3, 4, 5], led to the exciting idea that perhaps holography could be understood in a spacetime with a different geometry than Anti-de Sitter (AdS). This idea consequently led to a series of exploratory works on Schrödinger holography (see [7] for a recent review and references therein).

Holography, however, can only be properly tested when the bulk description and the boundary theory are simultaneously available. In the latter case - the dual field theory side - the authors of [3, 4, 5] argued that Schrödinger geometries (Sch) embedded in IIB string theory via the null Melvin twist should have a dual quantum field theory description in terms of a variant of Discrete Light Cone Quantization (DLCQ) (in the language of [4], this is referred to as  $\text{DLCQ}_\beta$  where  $\beta$  is the parameter describing the deformation from null AdS to Schrödinger), yielding a non-relativistic conformal field theory (CFT), e.g. of  $\mathcal{N} = 4$  Super Yang-Mills (SYM). However, the exact details and the precise prescription for performing this variant of DLCQ of SYM are still lacking, and it is therefore important to keep probing the properties of this dual field theory by means of well defined bulk calculations.

Thus taking a purely gravitational perspective, and noting that certain Schrödinger geometries can be obtained from AdS via the null Melvin twist [3, 4, 5, 8, 9, 10], one is naturally led to trying to understand

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<sup>1</sup>See also [6] for earlier work on Schrödinger spacetimes.

which geometrical and gravitational properties of AdS carry over to (or are inherited by) Schrödinger and which ones do not (this was also the point of view adopted e.g. previously in [11, 12, 13, 7]).

It has been argued in [3] that this null Melvin twist (or null dipole, or  $\text{DLCQ}_\beta$ ) deformation leaves invariant certain observables and correlation functions in the dual field theory, in particular those of local operators without external momenta. On the other hand, examples of observables widely studied in the context of AdS/CFT which do not fall into this class are expectation values of the Wilson loop operators [14, 15]. It is therefore of particular interest to study the potential gravitational duals of such Wilson loop (and Wilson surface) operators in the Schrödinger context, and this is the aim and content of this paper.

In the bulk theory on  $\text{AdS}_5 \times S^5$ , a Wilson loop is described by a configuration where a string is stretched into the bulk and whose endpoints, representing a quark-antiquark pair, are fixed on the AdS boundary. In particular, the free energy  $\mathcal{F}_{\text{loop}}$  of these configurations, obtained via a bulk calculation, has the interpretation of the quark-antiquark potential, which in this case, at zero temperature, is Coulomb-like and hence proportional to the inverse length  $L$  of the quark-antiquark pair [14, 15]. This dependence of the free energy on  $L$  is in accordance with the expected form for a conformally invariant dual theory. Therefore, Wilson loops can be used as probes of some of the properties of the dual field theory and hence it is interesting to search for an analogous statement in  $\text{Sch}_5 \times S^5$ , obtained via the null Melvin twist of the extremal D3-brane [3, 4, 5]. In fact, we will see that the free energy of these configurations is no longer Coulomb-like, but instead interpolates between a result proportional to the inverse length  $L$  and a result proportional to the inverse square of the length for large enough  $\ell = \beta R^2$  - the parameter describing the deformation from  $\text{AdS}_5$  with radius  $R$  into  $\text{Sch}_5$ . Similar arguments apply to the case of the expectation value of Wilson surface operators and their bulk duals defined in  $\text{AdS}_7 \times S^4$  [14], and correspondingly in  $\text{Sch}_7 \times S^4$  - their null Melvin twisted version [16].

Wilson loop operators, dual to strings with endpoints on the  $\text{AdS}_5 \times S^5$  boundary, have a supergravity description in terms of fully backreacted fundamental strings via the open/closed string duality. A similar example is that of point particles blown up into D3-branes orbiting a great circle of the  $S^5$ , the so called giant gravitons [17], which when fully backreacted have a supergravity description in terms of LLM geometries [18]. Therefore, constructing bulk configurations with holographic duals can be thought of as constructing supergravity solutions. An interesting question that arises is how these configurations are modified when the background is at finite temperature. In this context, finite temperature corrections to the quark-antiquark potential have been obtained in [19, 20] by solving the Nambu-Goto action, i.e. by placing an extremal, zero-temperature, probe on a black hole background. This method however does not take into account the thermal excitations of the string degrees of freedom, which in practice means that the probe is not in thermal equilibrium with the background. A way to take this into account in the strong coupling, supergravity regime,<sup>2</sup> is to use the blackfold approach [22, 23]. This method, which we briefly recall in Sec. 2.1 consists in placing probe black fundamental strings or black branes, hence with an intrinsic temperature, on a given background and requiring thermodynamic equilibrium [24, 25, 26, 27, 28, 29, 30, 31]. In the case of Wilson loops, this has been analysed in [26] and led to quantitative and qualitative new features which were not captured by the analysis of [19, 20]. Here, we describe the first application of this method to Schrödinger spacetimes and construct perturbatively supergravity black strings and black membranes. We also consider black membranes, dual to Wilson surface operators, in  $\text{AdS}_7 \times S^4$  using this black probe method and find new features which were not

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<sup>2</sup>See the recent analysis of [21] for the thermal DBI action at weak coupling for the D3-brane.

observed in [32].

This paper is organised as follows. In Sec. 2 we begin by reviewing the black probe method based on the blackfold approach for supergravity/M-theory [24, 33, 27] which will allow us to construct black string and black membrane geometries dual to certain gauge theory operators. Since the Schrödinger spacetimes that we consider are obtained by performing the null Melvin twist of known supergravity solutions and consequently taking their near horizon limit, the resulting geometry is still a solution of the complete set of supergravity equations and hence we can apply the black probe method developed for supergravity/M-theory [24, 33, 27]. In Sec. 2.2 we review the work of [26], where Wilson loops at finite temperature were analysed, and extend it to thermal AdS. In Sec. 3 we consider analogous configurations in  $\text{Sch}_5 \times S^5$  and find three distinct types of configurations. One of these interpolates between the AdS configuration and another Schrödinger configuration where the string has boundary endpoints along a spatial direction. Another of these configurations consists of a string with boundary endpoints separated along a null direction. We compare these configurations with the AdS case. In Sec. 4 we apply the method to bulk duals to Wilson surfaces in  $\text{AdS}_7 \times S^4$  and find dominant temperature corrections compared to the work of [32]. Afterwards we construct several such configurations in  $\text{Sch}_7 \times S^4$ . In Sec. 5 we discuss open problems and future research directions. App. A contains some details on the solution space of black membranes in  $\text{Sch}_7 \times S^4$ .

## 2 Black strings in Anti-de Sitter

In this section we begin by reviewing the black probe method based on the blackfold approach for supergravity and M-theory developed in [24, 33, 27] which allows us to perturbatively construct the black hole configurations presented throughout this work. We then review how this method was applied in order to construct black strings in  $\text{AdS}_5 \times S^5$  in the deconfined phase dual to finite temperature Wilson loops [26] and generalise this construction to the confined phase. The fact that the leading order temperature contribution to the quark-antiquark potential is due to the blackness of the probe, and not due to the characteristics of the background, is highlighted. We further notice that previous results based on the method of extremal, zero-temperature, probes [19, 20] are recovered by taking a non-trivial double scaling limit where the resulting finite temperature configuration is composed of locally flat, zero-temperature, fundamental strings appropriately redshifted. This will then serve as a means of comparison with the same configuration in Schrödinger spacetime which will be analysed in Sec. 3.

### 2.1 Black probe method

The open/closed string duality states that Dp-branes and fundamental  $F_1$  strings are described either by a low-energy effective action at weak coupling, such as the Dirac-Born-Infeld (DBI) action and the Nambu-Goto (NG) action, or by a fully backreacted supergravity solution at strong coupling. When the low-energy action is used in the context of AdS/CFT in order to find Dp-brane or  $F_1$  configurations in a closed string background, they can be related to certain operators in the dual zero-temperature gauge theory and simultaneously have a supergravity regime where the Dp-branes or  $F_1$  strings are fully backreacted. The clearest example of this is that of giant gravitons which are solutions of the DBI action coupled to the background Ramond-Ramond field [17] dual to large N gauge theory operators [34, 35] and when fully backreacted give rise to the LLM geometries [18].

If the closed string background has a finite temperature, the usual DBI or NG action is no longer a good description of the weak coupling dynamics. Instead, for small temperatures one should quantise the world volume theory and include the thermalisation of the brane/string degrees of freedom which heat up as the brane is immersed in a background at finite temperature. In general, the thermal version of the DBI action at weak coupling is not known but it has recently been derived for the case of the D3-brane [21]. On the other hand, finite temperature configurations at weak coupling will have a strong coupling, fully backreacted, regime where the corresponding objects will be non-extremal, finite-temperature, black hole geometries.

In this paper we are interested in the strong coupling regime  $g_s N_{(p)} \gg 1$  with  $g_s \ll 1$  being the string coupling and  $N_{(p)} \gg 1$  the number of probe Dp-branes (in the case  $p = 1$  we write  $N_{(1)} = k$  to designate the number of probe  $F_1$  strings), where we have a supergravity/M-theory description. In this case, the black hole solutions can be constructed order-by-order in a derivative expansion using the blackfold approach [22, 23, 24, 33, 27] and their construction is independent of there being a weakly coupled description of the same object or not. When considering AdS backgrounds we will be constructing the finite temperature dual of the gauge theory Wilson loops and Wilson surfaces operators but when considering Schrödinger backgrounds we take a more gravitational point of view and construct different black hole geometries, regardless of whether or not these configurations are dual to some gauge theory operators in a non-relativistic CFT.

The blackfold approach consists of placing a number of  $N_{(p)}$  black probe branes in a background spacetime with metric  $g_{\mu\nu}$  and wrapping them along a world volume  $\mathcal{W}_{p+1}$  with coordinates  $\sigma^a$ , whose position in the ambient space is parametrised by the mapping functions  $X^\mu(\sigma^a)$ . In terms of a metric, we are constructing solutions for which their near-horizon geometry is that of a supergravity/M-theory black brane and the far away asymptotics is given by the arbitrary metric  $g_{\mu\nu}$  via a matched asymptotic expansion procedure.<sup>3</sup> It has been shown that for this procedure to yield a valid gravity solution, in the case of neutral black branes it must satisfy, for stationary configurations, the equilibrium condition [36, 37]

$$T^{ab} K_{ab}{}^\rho = 0 \quad , \quad (2.1)$$

which, when ignoring possible background dilatonic fields and fluxes, can also be shown to hold for supergravity/M-theory charged branes [38]. In Eq. (2.1), we have introduced the effective stress-energy tensor  $T^{ab}$  describing the black probes and the extrinsic curvature of the geometry  $K_{ab}{}^\rho$  which can be obtained via the relation  $K_{ab}{}^\rho = \nabla_a \partial_b X^\rho$  where the covariant derivative  $\nabla_a$  is compatible with  $g_{\mu\nu}$  as well as with the induced metric  $\gamma_{ab} = g_{\mu\nu} \partial_a X^\mu \partial_b X^\nu$ . When the world volume has boundaries, Eq. (2.1) must be supplemented with boundary conditions, which in this case are [27]

$$T^{ab} \eta_b|_{\partial\mathcal{W}_{p+1}} = 0 \quad , \quad J^{ba_1 \dots a_p} \eta_b|_{\partial\mathcal{W}_{p+1}} = 0 \quad , \quad (2.2)$$

where  $\eta_b$  is a unit normalised orthogonal covector to the boundary  $\partial\mathcal{W}_{p+1}$  and  $J^{a_1 \dots a_{p+1}}$  denotes the  $(p+1)$ -dimensional effective current carrying the conserved total charge  $Q_p$  of the black probe satisfying  $\partial_a Q_p = 0$ .

For supergravity/M-theory probes the effective stress tensor takes the form of a perfect fluid [33],

$$T^{ab} = P \gamma^{ab} + (\epsilon + P) u^a u^b \quad , \quad \epsilon + P = \mathcal{T} s \quad , \quad (2.3)$$

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<sup>3</sup>For an example of the implementation of this procedure for neutral black branes in pure vacuum see [36].

while the effective current reads

$$J^{a_1 \dots a_{p+1}} = Q_p \varepsilon^{a_1 \dots a_{p+1}} , \quad (2.4)$$

where  $P, \epsilon, \mathcal{T}, s$  denote the local pressure, energy density, temperature and entropy density respectively while  $u^a$  is the normalised ( $u^a u_a = -1$ ) fluid velocity which, for stationary configurations, must be aligned with a world volume Killing vector field  $\mathbf{k}^a$  such that  $u^a = \mathbf{k}^a / |\mathbf{k}|$ , with  $|\mathbf{k}| = \sqrt{-\gamma_{ab} \mathbf{k}^a \mathbf{k}^b}^{\frac{1}{2}}$ . The indices  $a, b, c, \dots$  label the  $(p+1)$  world volume directions while the greek indices  $\mu, \nu, \dots$  span the entire spacetime dimensions and  $\varepsilon^{a_1 \dots a_{p+1}}$  is the Levi-Civita tensor on  $\mathcal{W}_{p+1}$ .

The fact that the probe is composed of a fluid with non-zero temperature gives rise to the required degrees of freedom that need to be heated up when placing the configuration in a background spacetime at finite temperature. The thermodynamic fluid variables, as well as the charge  $Q_p$  and corresponding chemical potential  $\Phi_p$ , are given in terms of the horizon radius  $r_0$  and the charge parameter  $\alpha$  of the black brane by [33]

$$\epsilon = \frac{\Omega_{(n+1)}}{16\pi G} r_0^n (n+1 + n \sinh^2 \alpha) , \quad P = -\frac{\Omega_{(n+1)}}{16\pi G} r_0^n (1 + n \sinh^2 \alpha) , \quad (2.5)$$

$$\mathcal{T} = \frac{n}{4\pi r_0 \cosh \alpha} , \quad s = \frac{\Omega_{(n+1)}}{4G} r_0^{n+1} \cosh \alpha , \quad (2.6)$$

and

$$Q_p = \frac{\Omega_{(n+1)}}{16\pi G} n r_0^n \sinh \alpha \cosh \alpha , \quad \Phi_p = \tanh \alpha . \quad (2.7)$$

Here, we have focused on  $D = 10$  and  $D = 11$  supergravity/M-theory branes and introduced Newton's constant  $G$  and the volume of an  $(n+1)$ -sphere  $\Omega_{(n+1)}$  where the dimension  $n$  is defined via  $D = n + p + 3$ .

In the case of stationary configurations the equation of motion (2.1) can be integrated to an action [33]

$$I[X^\mu(\sigma^a)] = \int_{\mathcal{W}_{p+1}} d^{p+1} \sigma \sqrt{-\gamma} P , \quad (2.8)$$

where  $\gamma$  is the determinant of the induced metric. The branes characterised by these thermodynamic quantities admit an extremal limit where the local temperature  $\mathcal{T} \rightarrow 0$ . This limit is obtained by sending  $\alpha \rightarrow \infty$  such that  $\mathcal{T} \rightarrow 0$  and  $r_0 \rightarrow 0$  while keeping the total charge  $Q_p$  constant. It is easy to see that in this case  $P \rightarrow -Q_p$ , i.e.,

$$\alpha \rightarrow \infty , \quad r_0 \rightarrow 0 , \quad P \rightarrow -Q_p . \quad (2.9)$$

In the context of AdS/CFT where the charge is quantised such that  $Q_p = N_{(p)} T_{(p)}$  with  $T_{(p)}$  being the tension of the Dp-brane (or in the case  $p = 1$ ,  $T_{(1)} = T_{F_1}$  being the tension of the fundamental  $F_1$  string), the action (2.8) yields the zero-temperature extremal action

$$I[X^\mu(\sigma^a)]_{\text{ext}} = -N_{(p)} T_{(p)} \int_{\mathcal{W}_{p+1}} d^{p+1} \sigma \sqrt{-\gamma} , \quad (2.10)$$

which differs from the weak coupling DBI description, valid for  $g_s \ll 1$  and  $N_{(p)} = 1$ , by a multiplicative factor of  $N_{(p)}$ . The fact that we have obtained  $N_{(p)}$  times the DBI action reflects the supersymmetry of the solutions.<sup>4</sup>

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<sup>4</sup>For supersymmetric deformations of blackfold geometries see [39].

Since solutions of (2.8) are time independent, the problem can be posed in an Euclidean version, where the action (2.8) is Wick rotated such that  $\sigma^0 = \tau \rightarrow i\tau$  and integrated over the time circle of radius  $1/T$  where  $T$  is the background temperature related to the local brane temperature via

$$T = \mathbf{k} \mathcal{T} . \quad (2.11)$$

This relation is one of the main improvements of the black probe method compared to the extremal probe method as it expresses that the probe temperature  $\mathcal{T}$  must be in equilibrium with the background temperature  $T$  via a redshift factor  $\mathbf{k}$ . The resulting quantity, after the Wick rotation, is the Gibbs free energy

$$\mathcal{F}[X^\mu(\sigma^a)] = - \int_{\mathcal{B}_p} dV_{(p)} R_0 P , \quad (2.12)$$

where  $dV_{(p)}$  is the volume form on the  $p$ -spatial world volume  $\mathcal{B}_p$  and  $R_0$  is the local redshift factor obtained by computing the norm of a timelike background Killing vector  $\xi$  on the world volume, i.e.,  $R_0^2 = -\xi^\mu \xi_\mu|_{\mathcal{W}_{p+1}}$ . The entropy  $S$  and the energy/mass  $M$  can be obtained from (2.12) via the relations

$$S = -\frac{\partial \mathcal{F}}{\partial T} , \quad M = \mathcal{F} - TS , \quad (2.13)$$

where we have used the first law of thermodynamics  $d\mathcal{F} = dM - TdS$  and assumed the fluid to be static, as it will be such cases that we will deal with in this paper. In order to obtain different black hole configurations we will be varying the free energy (2.12) and solving the equations of motion (2.1) for different configurations of the scalars  $X^\mu(\sigma^a)$  in several ambient spacetimes.

## 2.2 Black strings dual to finite temperature Wilson loops

In this section we review the finite temperature Wilson loops in the deconfined phase constructed in [26] by wrapping black strings using the method described above and extend it to the confined phase. The construction in the confined phase, even though of less physical interest since it does not exhibit a phase transition and hence no Debye screening effect, presents itself as a more striking example which highlights the differences between the black probe method and the extremal probe method [19, 20]. We also show how the work of [19, 20] is recovered by taking a non-trivial double scaling limit.

We consider the AdS black hole background with AdS radius  $R$  and metric

$$ds^2 = \frac{R^2}{z^2} (-f(z)dt^2 + f(z)^{-1}dz^2 + dx_i^2) + R^2 d\Omega_{(5)}^2 , \quad f(z) = 1 - \gamma \frac{z^4}{z_0^4} , \quad (2.14)$$

where  $i = 1, 2, 3$ ,  $d\Omega_{(5)}^2$  is the metric on the five-sphere and  $z_0 = 1/(\pi T)$  is the location of the horizon and  $T$  its temperature. The boundary is located at  $z = 0$ . The parameter  $\gamma$  controls whether we are in the deconfined phase ( $\gamma = 1$ ) or in the confined phase ( $\gamma = 0$ ). In the latter case, the metric corresponds to thermal AdS. The Hawking-Page transition occurs when the background temperature exceeds  $T_{\text{HP}} \sim 1/R$  for which the case  $\gamma = 1$  is thermodynamically preferred. We have ignored the five-form fluxes present in the background since they do not play a role in the examples we consider.

## Embedding and solution

We now follow the prescription of [26] and choose a static string embedding ( $p = 1$ ,  $n = 6$ ) with its endpoints on a boundary spatial direction separated by a length  $L$ , at the locations  $x_1 = 0$  and  $x_1 = L$ , and stretched into the bulk along the  $z$ -direction. The embedding map is

$$t = \tau \quad , \quad z = \sigma \quad , \quad x_1 = x(\sigma) \quad , \quad x_2 = x_3 = d\Omega_{(5)} = 0 \quad . \quad (2.15)$$

With this choice the problem becomes symmetric around  $x_1 = L/2$ . The string is stretched from  $(x_1, z) = (0, 0)$  to  $(x_1, z) = (L/2, \sigma_0)$  and back again to  $(x_1, z) = (L, 0)$ . The induced metric on  $\mathcal{W}_2$  becomes

$$\gamma_{ab} d\sigma^a d\sigma^b = \frac{R^2}{\sigma^2} (-f(\sigma) d\tau^2 + (x'(\sigma)^2 + f(\sigma)^{-1}) d\sigma^2) \quad , \quad f(\sigma) = 1 - \gamma(\pi T \sigma)^4 \quad , \quad (2.16)$$

where the prime denotes a derivative with respect to  $\sigma$ . The norm of the pull-back of the timelike Killing vector field  $\partial_t$  onto the world volume yields the redshift factor  $R_0 = (R/\sigma)\sqrt{f}$  and, since the string is static, we have that  $\mathbf{k} = R_0$ .

Evaluating the determinant of (2.16) and introducing it in (2.12) leads to the free energy

$$\mathcal{F}[x(\sigma)] = -2A \left( \frac{3}{2\pi T} \right)^6 \int_0^{\sigma_0} d\sigma \sqrt{1 + f(\sigma) x'(\sigma)^2} G(\sigma) \quad , \quad G(\sigma) = \frac{R^8}{\sigma^8} f(\sigma)^3 \frac{1 + 6 \sinh^2 \alpha}{\cosh^6 \alpha} \quad , \quad (2.17)$$

where we have defined  $A = \Omega_{(7)}/(16\pi G)$  and used Eqs. (2.5)-(2.7) as well as (2.11). The overall factor of 2 accounts for the fact that we are gluing two symmetric patches of string around  $x_1 = L/2$ . The equation of motion that follows from varying (2.17) is

$$\left( \frac{f(\sigma) x'(\sigma)}{\sqrt{1 + f(\sigma) x'(\sigma)^2}} G(\sigma) \right)' = 0 \quad . \quad (2.18)$$

We consider solutions for which  $x'(\sigma) > 0$  and which satisfy the boundary conditions  $x(0) = 0$  and  $x'(\sigma) \rightarrow \infty$  for  $\sigma \rightarrow \sigma_0$ . This yields the solution

$$x'(\sigma) = \left( \frac{f(\sigma)^2 G(\sigma)^2}{f(\sigma_0) G(\sigma_0)^2} - f(\sigma) \right)^{-\frac{1}{2}} \quad . \quad (2.19)$$

In the case  $\gamma = 1$ , this agrees with the result found in [26] while in the case  $\gamma = 0$ , and hence  $f(\sigma) \rightarrow 1$ , we obtain

$$x'(\sigma)|_{\gamma=0} = \left( \frac{G(\sigma)^2}{G(\sigma_0)^2} - 1 \right)^{-\frac{1}{2}} \quad , \quad G(\sigma)|_{\gamma=0} = \frac{R^8}{\sigma^8} \frac{1 + 6 \sinh^2 \alpha}{\cosh^6 \alpha} \quad . \quad (2.20)$$

We further note that the boundary conditions (2.2) are ensured with the choices we made. In order to see this explicitly we note that the world volume of the string  $\mathcal{W}_2$  composed of the two symmetric patches only has a boundary at  $\sigma = 0$ . The unit normalised, orthogonal covector to the world volume boundary is

$$\eta_b d\sigma^b = \frac{R}{\sigma} \frac{\sqrt{1 + f(\sigma) x'(\sigma)^2}}{\sqrt{f(\sigma)}} d\sigma \quad . \quad (2.21)$$

With this we evaluate (2.2) leading to

$$T^{\sigma\sigma} \eta_\sigma|_{\sigma=0} = -Q_1 \frac{(1 + 6 \sinh^2 \alpha)}{6 \sinh \alpha \cosh \alpha} \frac{\sigma}{R} = 0 \quad , \quad J^{\tau\sigma} \eta_\sigma|_{\sigma=0} = Q_1 \frac{\sigma}{R} = 0 \quad , \quad (2.22)$$



where we have used Eqs. (2.5)-(2.7) and the fact that  $f(\sigma) \rightarrow 1$  and  $G(\sigma) \rightarrow \infty$  when  $\sigma \rightarrow 0$ .

Finally, we note that the equation of motion (2.18) also admits the solution  $x'(\sigma) = 0$  corresponding to two straight strings stretched from the boundary to the black hole horizon ( $\sigma = z_0$ ) in the case  $\gamma = 1$  or to the origin of thermal AdS ( $\sigma = \infty$ ) in the case  $\gamma = 0$ . In terms of the dual gauge theory, this represents the Polyakov loop. In this case, the geometry also satisfies the boundary conditions (2.2), in particular, at  $\sigma = 0$  we obtain the same result as in (2.22) valid for any choice of  $\gamma$ .

## Solution space

Here we wish to analyse the allowed region of solution space. As in [26], we first introduce a dimensionless coordinate  $\hat{\sigma} = \pi T \sigma$  and define a parameter  $\kappa$  using the Eqs.(2.7) and (2.11) according to

$$\kappa \equiv \frac{2^5 Q_1}{3^7 A R^6} = \frac{f(\hat{\sigma})^3}{\hat{\sigma}^6} \frac{\sinh \alpha(\hat{\sigma})}{\cosh^5 \alpha(\hat{\sigma})} , \quad (2.23)$$

where  $f(\hat{\sigma}) = 1 - \gamma \hat{\sigma}^4$ . The parameter  $\kappa$  as we will see below distinguishes between the extremal probe  $\kappa = 0$  and the non-extremal probe  $\kappa \neq 0$ .<sup>5</sup> Therefore we seek to understand the effect of corrections proportional to  $\kappa$  in these configurations. Because of the fact that the ratio  $\sinh \alpha / \cosh^5 \alpha$  is bounded from above by the value  $2^4/5^{5/2}$ , the black string cannot be stretched all the way to the horizon for any value of  $\kappa$ . The maximum distance into the bulk it can attain is given by the critical distance

$$\hat{\sigma}_c^2 = \frac{1}{\gamma} \left( \sqrt{\gamma + \frac{5^{5/3}}{2^{14/3}} \kappa^{2/3}} - \frac{5^{5/6}}{2^{7/3}} \kappa^{1/3} \right) . \quad (2.24)$$

In the case  $\gamma = 1$  this agrees with the result obtained in [26] and attains its maximum value  $\hat{\sigma}_c = 1$ , where it touches the horizon, when  $\kappa = 0$  and decreases for increasing  $\kappa$  meaning that when  $\kappa > 0$  it cannot reach all the way to the horizon. In the confined phase for which  $\gamma = 0$  we instead get

$$\hat{\sigma}_c^2|_{\gamma=0} = \frac{2^{4/3}}{5^{5/6}} \kappa^{-\frac{1}{3}} , \quad (2.25)$$

which means that for  $\kappa = 0$  the string can stretch all the way to the origin of thermal AdS but for  $\kappa > 0$  it cannot be stretched more than a finite amount. The fact that the distance (2.24) exists is due to the fact that the local temperature of the black probe  $\mathcal{T}$  has a maximum value for a given value of the charge  $Q_1$ . By looking at (2.24) one may be tempted to say that in the case  $\gamma = 0$  the string can be stretched more into the bulk than when  $\gamma = 1$ . This however is only apparent as by comparing the invariant spacetime lengths along the  $z$ -direction one finds that

$$\hat{\sigma}_c|_{\gamma=0} = \frac{\hat{\sigma}_c|_{\gamma=1}}{f(\hat{\sigma}_c|_{\gamma=1})^{\frac{1}{2}}} , \quad (2.26)$$

and therefore the strings can be stretched exactly the same distance in both cases for a given value of  $\kappa$ .

We now comment on how the different regions of parameter space are connected to each other and how the work of [14, 15, 19, 20] is recovered. For a given configuration with finite  $T, \kappa$ , as indicated in Fig. 1, there are two ways of making the black probe extremal. According to (2.11) we have that

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<sup>5</sup>This distinction only makes sense if the charge  $Q_1$  is non-vanishing. If the charge vanishes there is no extremal limit since uncharged supergravity/M-theory black branes do not admit an extremal limit. Around Eq. (2.27) we will explain how the extremal limit  $\kappa = 0$  with  $Q_1$  non-vanishing can be obtained.

- $T \rightarrow 0$  and  $\mathcal{T} \rightarrow 0$  with  $\mathbf{k}\mathcal{T} \rightarrow 0$ .
- $\mathcal{T} \rightarrow 0$  and  $\mathbf{k}\mathcal{T}$  remains non-zero in such a way that  $T$  remains finite and non-zero.

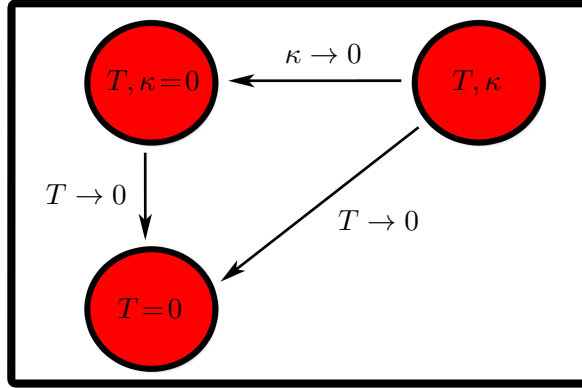


Figure 1: Diagram depicting the solution space and its limits. The solutions at  $T = 0$  for  $\kappa = 0$  and  $\kappa \neq 0$  are equivalent. This diagram is generic for all the configurations constructed in this paper.

In the first case the background temperature is set to zero and we recover the solution at zero-temperature found in [14, 15] regardless of the value of  $\kappa$ . In the second case, the black brane becomes extremal and is appropriately locally redshifted such that  $\mathbf{k}\mathcal{T}$  remains non-zero. This can only happen if we send  $\kappa \rightarrow 0$  for finite  $Q_1$ , according to Eq. (2.23).<sup>6</sup> This is in fact the first example of a black hole solution where the near-horizon geometry is locally composed of extremal, zero-temperature, black branes but for which the global geometry has a finite temperature  $T$ . In this case the results of [19, 20] are reproduced. If we would now turn off the temperature we would find the results of [14, 15]. In more physical terms, we can rewrite  $\kappa$  using the AdS/CFT dictionary for the radius  $R^4 = \lambda l_s^4$ , the string coupling  $4\pi g_s = \lambda/N$  as well as the relation for the fundamental string tension  $T_{F1} = (2\pi l_s^2)^{-1}$ , where  $\lambda$  is the 't Hooft coupling of the  $\mathcal{N} = 4$  SYM theory,  $N$  is the number of black D3-branes whose near-horizon limit gave rise to (2.14) and  $l_s$  is the string length. Newton's constant can be written in terms of the Planck length  $l_p$  according to  $16\pi G = (2\pi)^{D-3} l_p^{D-2}$ . In type IIB supergravity we have that  $l_p^8 = g_s^2 l_s^8$ . Using this, the parameter  $\kappa$  can be written as [26]

$$\kappa = \frac{2^7}{3^6} \frac{k \sqrt{\lambda}}{N^2} . \quad (2.27)$$

Therefore, we see that the work of [19, 20] is recovered in the strict limit  $N \rightarrow \infty$  while still satisfying local thermodynamic equilibrium with the black probe in accordance with Eq. (2.11).<sup>7</sup>

There is also another region of solution space which has not been previously explored, though of less holographic interest, as it does not connect with the part of solution space scanned by the extremal probe. This region is attained by sending  $\kappa \rightarrow 0$  while keeping  $T \neq 0$  but sending  $Q_1 \rightarrow 0$ . This branch of solutions does not admit an extremal limit since uncharged supergravity/M-theory black branes do not admit an extremal limit. However, this region is beyond the regime of validity of our method, as we will see below.

<sup>6</sup>As it will be explained below, this implies taking the strict limit  $N \rightarrow \infty$ .

<sup>7</sup>The relation  $R^4 = \lambda l_s^4$  implies that if we send  $\kappa \rightarrow 0$  while keeping  $Q_1$  non-zero then we must have that  $N \rightarrow \infty$  such that  $\sqrt{\lambda}/N^2 \rightarrow 0$ .

## Regime of validity

We now wish to consider the regime of validity of the solutions. We consider the regime connected to the extremal probe ( $\alpha \rightarrow \infty$ ) for which the characteristic length of the black fundamental string is its charge radius  $r_c = r_0 \sinh^2 \alpha$ .

- For the solutions to be valid, one needs to require that quantum corrections are sub-leading, i.e.,  $r_c \gg l_s$  which implies that  $kg_s^2 \gg 1$ . Furthermore, for the near-horizon geometry to be considered locally flat one must require  $r_c \ll R$  implying that  $k \ll N$ , given that  $g_s \ll 1$ . With this one concludes that we must require,

$$1 \ll k \ll N \quad . \quad (2.28)$$

The charge radius can also be written in terms of the parameter  $\kappa$ , which when using (2.23) reads  $r_c = R\kappa^{1/6}$ . Therefore, requiring the near-horizon geometry to be locally flat when compared to the two background scales  $r_c \ll R$  and  $r_c \ll 1/T$  one obtains

$$\kappa \ll 1 \quad , \quad RT \ll \kappa^{-\frac{1}{6}} \quad . \quad (2.29)$$

The latter condition, imposing a very loose upper bound on the temperature of the black hole, is only present in the case  $\gamma = 1$ . In terms of the 't Hooft coupling the former condition simply implies that  $\lambda \ll N^2$ . These conditions had already been derived in [26].

- One needs to further require that variations in the local temperature along the string are small compared to the background temperature, i.e.,

$$r_c \frac{\mathcal{T}'}{T} = \frac{r_c}{R\sqrt{f}} + 2\gamma \frac{(\pi T \sigma)^4}{R f^{3/2}} \ll 1 \quad . \quad (2.30)$$

In the case  $\gamma = 0$  this condition reduces to  $r_c \ll R$ , which has already been ensured by taking  $\kappa \ll 1$ . In the case  $\gamma = 1$  this is also satisfied when the distances are such that  $\sigma \ll z_0$ , otherwise, for small  $\kappa$  and for distances  $\sigma \sim z_0$  one finds  $z_0 - \sigma \gg \kappa^{1/9} z_0$ . In particular, this implies that for  $\gamma = 1$  the approximation breaks down for distances close to  $\sigma \sim \sigma_c$ .

- One should also require that the black string is thin compared to the local extrinsic length scale. For this one evaluates the scale of the extrinsic curvature via<sup>8</sup>

$$L_{\text{ext}}(\sigma) = |K^\rho N_\rho|^{-1} = \frac{R}{\sigma} \frac{(1 + f x'(\sigma)^2)^{3/2}}{f |x''(\sigma)|} \quad , \quad (2.31)$$

where  $K^\rho = \gamma^{ab} K_{ab}{}^\rho$  is the mean extrinsic curvature and  $N_\rho$  is a unit normalised covector orthogonal to  $\mathcal{W}_2$ . The length scale  $L_{\text{ext}}(\sigma)$  is minimal when  $\sigma = \sigma_0$ . In the confined phase we find that for any  $\sigma_0$ , including  $\sigma_c$ , imposing  $r_c \ll L_{\text{ext}}(\sigma)|_{\gamma=0}$  implies again  $r_c \ll R$ , which has already been ensured. In the deconfined phase, for distances  $\sigma \ll z_0$  one again finds  $r_c \ll R$  but for distances  $\sigma \sim z_0$  one finds that  $L_{\text{ext}}(\sigma_c) \sim r_c$  and hence the approximation breaks down. One should then require  $z_0 - \sigma \gg \kappa^{1/9} z_0$  as above.<sup>9</sup>

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<sup>8</sup>In evaluating (2.31) we have ignored derivatives of the metric since we have taken into account the variations of the background. One can explicitly check that the regimes of validity do not change if one considers these variations in (2.31).

<sup>9</sup>In general, besides the length scale of variations of the extrinsic curvature (2.31), one should also consider the length scale of variations associated with the intrinsic curvature of the world volume (2.16). It is a straightforward exercise to check that the intrinsic curvature scale is of the same order as (2.31).

We note that in the confined phase, in contrast with the deconfined phase, the approximation does not break down near  $\sigma \sim \sigma_c$  and is valid all the way up to  $\sigma_c$ . The reason for this is simple. In the deconfined phase, the presence of the black hole introduces large gradients in the curvature near its horizon which induces large gradients in the extrinsic curvature and local temperature of the black string. On the other hand, in the confined phase, there is no horizon and large gradients are not present as one approaches the centre of AdS.

The straight string with  $x'(\sigma) = 0$  leads to a divergent  $L_{ext}(\sigma)$  and hence also satisfies the validity requirements. In the confined phase, the Polyakov loop can be extended all the way to  $\sigma_c$  but it cannot end there as it would not satisfy the boundary conditions (2.2). In the deconfined phase the approximation breaks down near  $\sigma_c$  but if it did not then the string could at most extend up to  $\sigma_c$ , which would again not satisfy the boundary conditions. We therefore conclude that, the Polyakov loop at finite temperature cannot be constructed using this method.

Finally, we comment on the branch of solutions connecting to uncharged black strings ( $\alpha = 0$ ). In this case the relevant scale characterising the black string is the energy density radius  $r_\epsilon \sim r_0$ . Using Eqs. (2.5)-(2.7) and imposing  $r_\epsilon \ll R$  one finds that one should have

$$\frac{1}{\sigma T} \sqrt{f(\sigma)} \ll 1 \quad . \quad (2.32)$$

However, for finite  $T$  as it is required for this branch of solutions, one sees that near the boundary  $\sigma \sim 0$  the radius  $r_\epsilon$  diverges and the condition above cannot be satisfied. We thus conclude that this branch of solutions is beyond the scope of our method.

### Regularized free energy and large $N$ expansion

We now take a closer look at the free energy (2.17) and deduce some analytic results in a large  $N$  expansion. As noted in [26], the free energy (2.17) can be rewritten, using Eqs. (2.5)-(2.7) and the AdS/CFT dictionary, as

$$\mathcal{F}[x(\hat{\sigma})] = \sqrt{\lambda} kT \int_0^{\hat{\sigma}_0} \frac{d\hat{\sigma}}{\hat{\sigma}^2} (1 - X) \sqrt{1 + f x'(\hat{\sigma})^2} \quad , \quad X = 1 - \tanh \alpha - \frac{1}{6 \sinh \alpha \cosh \alpha} \quad , \quad (2.33)$$

and diverges at  $\sigma = 0$ . This divergence at  $\sigma = 0$  corresponds to an ultraviolet divergence in the gauge theory and can be dealt with by subtracting a piece to the free energy with the interpretation of being the mass of the W-boson in the gauge theory. The prescription developed in [26] consists of introducing the infrared cut-off  $\hat{\sigma}_{\text{cut}}$  and subtracting the piece

$$\mathcal{F}_{\text{sub}} = \sqrt{\lambda} kT \int_0^{\hat{\sigma}_{\text{cut}}} \frac{d\hat{\sigma}}{\hat{\sigma}^2} (1 - X) \quad , \quad (2.34)$$

such that the free energy difference  $\Delta\mathcal{F}$  can be written as

$$\Delta\mathcal{F} = \mathcal{F} - \mathcal{F}_{\text{sub}} = \mathcal{F}_{\text{loop}} - 2\mathcal{F}_{\text{P}} \quad , \quad (2.35)$$

where

$$\mathcal{F}_{\text{loop}} = \sqrt{\lambda} kT \left( -\frac{1}{\hat{\sigma}_0} + \int_0^{\hat{\sigma}_0} \frac{d\hat{\sigma}}{\hat{\sigma}^2} \left( (1 - X) \sqrt{1 + f x'(\sigma)^2} - 1 \right) \right) \quad , \quad (2.36)$$

is the regularised free energy associated with the Wilson loop configuration and

$$\mathcal{F}_P = -\frac{1}{2}\sqrt{\lambda}kT \left( \frac{1}{\hat{\sigma}_{\text{cut}}} + \int_0^{\hat{\sigma}_{\text{cut}}} \frac{d\hat{\sigma}}{\hat{\sigma}^2} X \right) , \quad (2.37)$$

is the regularised free energy associated with the Polyakov loop. One then needs a prescription for fixing  $\hat{\sigma}_{\text{cut}}$ , which can be done by letting  $\hat{\sigma}_{\text{cut}}$  asymptote to  $\hat{\sigma}_c$ . However, one should pay attention to the following facts. As explained previously, Polyakov loops with  $\kappa \neq 0$  are outside the regime of validity of this approach so one cannot trust the expression (2.37). On the other hand, in the strict case  $\kappa = 0$  the expression (2.37) can be trusted since the validity requirements are satisfied. In this case one can expand the function  $X$  in powers of  $\kappa$  ending up with a result that vanishes when  $\kappa \rightarrow 0$ . In the deconfined phase we have that  $\hat{\sigma}_c = 1$  when  $\kappa = 0$  and in the confined phase  $\hat{\sigma}_c \rightarrow \infty$  when  $\kappa = 0$ , therefore one finds in general that for  $\kappa = 0$  one has

$$\mathcal{F}_P = -\frac{1}{2}\gamma\sqrt{\lambda}kT , \quad (2.38)$$

which in the case  $\gamma = 0$  is in agreement with [14] and in the case  $\gamma = 1$  in agreement with [26]. One can then argue that this result is valid for all  $\kappa \ll 1$  [26]. Assuming that the free energy is linear in  $T$  for any  $\kappa$  (this is an assumption made implicitly in [26]), the free energy (2.37) must be of the form  $\mathcal{F}_P = \sqrt{\lambda}kTa(\kappa)$  for some function  $a(\kappa)$ . Therefore, using Eq. (2.13) one finds the entropy  $S_P = -\sqrt{\lambda}ka(\kappa)$  which is independent of  $T$  since  $\kappa$  is independent of  $T$ . However, at zero-temperature the solution obtained with an extremal probe ( $\kappa = 0$ ) and with a non-extremal probe ( $\kappa \neq 0$ ) must be equivalent to each other, as illustrated in the diagram of Fig. 1. Therefore we see that we must have in general  $a(\kappa) = a(0) = \gamma$  and hence that (2.38) is valid for all  $\kappa \ll 1$ .<sup>10</sup>

We now take a look at (2.36). Since solutions are valid for  $\kappa \ll 1$  we can analytically expand (2.36) in powers of  $\kappa$ . An expansion in  $\kappa$  means an expansion in large  $N$  according to (2.27) but also an expansion around  $\mathcal{T} = 0$ , meaning that an expansion in inverse powers of  $N$  necessarily requires turning on the thermal degrees of freedom of the black probe. In order to proceed further we note that an expansion around  $\mathcal{T} = 0$  can be recast as an expansion around  $\alpha \rightarrow \infty$ . Introducing  $\phi = \cosh^2 \alpha$  in (2.23) and expanding it around  $\phi = \infty$  one finds

$$\frac{\hat{\sigma}^6}{f(\hat{\sigma})^3} \kappa = \frac{1}{\phi^2} - \frac{1}{2\phi^3} - \frac{1}{8\phi^4} + \mathcal{O}(\phi^{-5}) . \quad (2.39)$$

Solving this for  $\phi$  and expanding it for small  $\kappa$  one obtains

$$\cosh^2 \alpha = \frac{f(\hat{\sigma})^{3/2}}{\hat{\sigma}^3 \sqrt{\kappa}} - \frac{1}{4} - \frac{5}{32} \frac{\hat{\sigma}^3 \sqrt{\kappa}}{f(\hat{\sigma})^{3/2}} + \mathcal{O}(\kappa) . \quad (2.40)$$

Using this expansion one can obtain an analytic result for the distance  $L$  between the endpoints of the black string for small  $\hat{\sigma}_0$ , which in the dual gauge theory is interpreted as the length of the quark-antiquark

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<sup>10</sup>However, in general one could consider finite temperature corrections in small  $T$  of the form  $k\sqrt{\lambda}T(RT)^m b_m(\kappa)$  with  $m > 0$ ,  $b_m(0) = 0$  and  $\kappa \ll 1$  (since small  $T$  means that the probe is near extremal). In order to accommodate these possible corrections and obtain the result (2.37), one must require that the  $(m+1)$ 'th derivative for all  $m$  with respect to  $T$  of (2.37) exhibits the same behaviour in the limit  $T \rightarrow 0$  for all  $\kappa$ , implying that  $b_m(\kappa) = 0$ . The Polyakov loop at finite temperature has a supergravity description in terms of two straight black strings. We see that this requirement implies that the black string entropy is independent of  $T$  for small temperatures, which may be a very strict requirement. Even though we use this type of argument throughout this paper, we note that none of the results obtained here are affected by this except the analysis of possible phase transitions from the Wilson loops/surfaces to Polyakov loops/surfaces.

pair,

$$LT = \frac{2}{\pi} \int_0^{\hat{\sigma}_0} \frac{\partial x(\hat{\sigma})}{\partial \hat{\sigma}} = \frac{2\sqrt{2\pi}}{\Gamma(\frac{1}{4})^2} \hat{\sigma}_0 + \left( \frac{\sqrt{2\pi}}{3\Gamma(\frac{1}{4})^2} - \frac{1}{6} \right) \sqrt{\kappa} \hat{\sigma}_0^4 - \gamma \frac{2\sqrt{2\pi}}{5\Gamma(\frac{1}{4})^2} \hat{\sigma}_0^5 + \mathcal{O}(\hat{\sigma}_0^7) + \mathcal{O}(\kappa) \quad . \quad (2.41)$$

From this expression we see that for small distances  $LT$  one needs a small bulk depth  $\hat{\sigma}_0$ . Furthermore, one sees that the effect of the black hole background only enters at order  $\hat{\sigma}_0^5$  and hence the leading order correction to  $LT$  comes from the thermal degrees of freedom of the black probe. One can perform a similar expansion in  $\mathcal{F}_{\text{loop}}$  and find

$$\mathcal{F}_{\text{loop}} = -\sqrt{\lambda} k T \left( \frac{\sqrt{\pi} \Gamma(\frac{3}{4})}{\Gamma(\frac{1}{4}) \hat{\sigma}_0} - \frac{\pi \Gamma(\frac{1}{4}) + \sqrt{\pi} \Gamma(\frac{3}{4})}{6\Gamma(\frac{1}{4})} \hat{\sigma}_0^2 \sqrt{\kappa} + \frac{2\sqrt{\pi} \Gamma(\frac{3}{4})}{\Gamma(\frac{1}{4})} \gamma \hat{\sigma}_0^3 + \mathcal{O}(\kappa) + \mathcal{O}(\hat{\sigma}_0^4) \right) \quad . \quad (2.42)$$

Again we see here that the leading order correction comes from the thermal degrees of freedom and not from the background. We now invert (2.41) and introduce it in (2.42) in order to find

$$\mathcal{F}_{\text{loop}} = -\frac{\sqrt{\lambda} k}{L} \left( \frac{4\pi^2}{\Gamma(\frac{1}{4})^4} + \frac{\Gamma(\frac{1}{4})^4}{96} \sqrt{\kappa} (LT)^3 + \gamma \frac{3\Gamma(\frac{1}{4})^4}{160} (LT)^4 + \dots \right) \quad . \quad (2.43)$$

First we note that the combination  $\mathcal{F}_{\text{loop}} L$  only depends on the scale invariant combination  $LT$ , exhibiting the conformal symmetry of the dual gauge theory. Secondly, the leading order correction in  $LT$  is due to the properties of the probe and not of the background. This means that this correction term is also present in thermal AdS. A comparison between the last two terms present in (2.43) tells us that the term proportional to  $\gamma$  will be sub-leading if [26]

$$LT|_{\gamma=1} \ll \frac{\sqrt{k} \lambda^{1/4}}{N} \quad , \quad (2.44)$$

which requires small enough temperatures. One also sees that in the strict limit  $N \rightarrow \infty$  the correction due to the background is dominant. From (2.43) we can also compute the black string entropy and mass using (2.13),

$$S_{\text{loop}} = \frac{\sqrt{\lambda} k}{TL} \left( \frac{\Gamma(\frac{1}{4})^4}{32} \sqrt{\kappa} (LT)^3 + \gamma \frac{3\Gamma(\frac{1}{4})^4}{40} (LT)^4 + \dots \right) \quad , \quad (2.45)$$

$$M_{\text{loop}} = -\frac{\sqrt{\lambda} k}{L} \left( \frac{4\pi^2}{\Gamma(\frac{1}{4})^4} + \frac{\Gamma(\frac{1}{4})^4}{24} \sqrt{\kappa} (LT)^3 + \gamma \frac{3\Gamma(\frac{1}{4})^4}{32} (LT)^4 + \dots \right) \quad . \quad (2.46)$$

### Comparison between phases

We now compare the distance between the quark-antiquark pair and the free energy in both phases. We begin by solving for the distance  $LT$  numerically for small values of  $\kappa$ . This is shown in the figures below. The dashed black line represents the value of  $LT$  for  $\kappa = 0$ . This can be obtained exactly analytically

$$LT|_{\kappa=0} = \frac{2\sqrt{2\pi}}{\Gamma(\frac{1}{4})^2} \hat{\sigma}_0 \sqrt{1 - \gamma \hat{\sigma}_0^4} {}_2F_1 \left( \frac{1}{2}, \frac{3}{4}, \frac{5}{4}; \gamma \hat{\sigma}_0^4 \right) \quad . \quad (2.47)$$

From the figures below one can see that in both phases, there is a critical distance  $\hat{\sigma}_c$  beyond which the solution terminates. In the confined phase, this difference is more pronounced when compared to the

$\kappa = 0$  case since in this case the extremal probe method tells us that the black string can extend all the way to the origin of AdS.

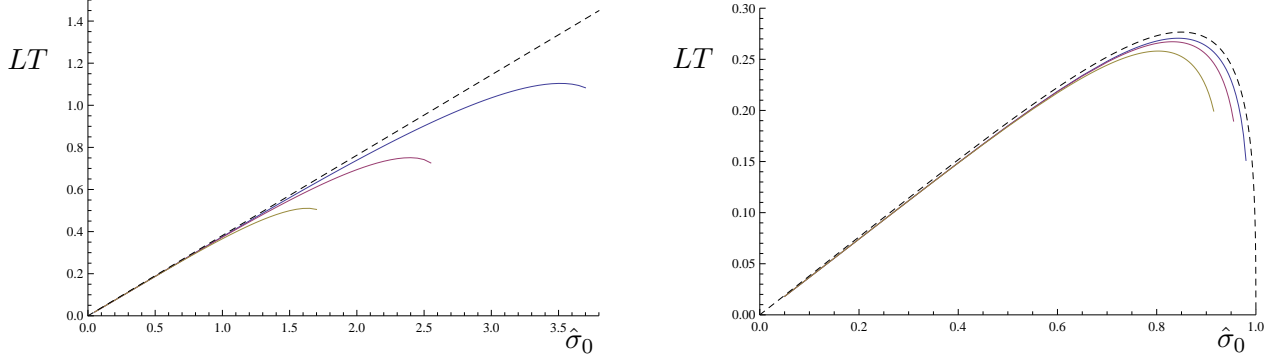


Figure 2: On the left is the plot of the distance  $LT$  as a function of  $\hat{\sigma}_0$  for  $\kappa = 0.01$  (yellow line),  $\kappa = 0.001$  (red line) and  $\kappa = 0.0001$  (blue line) in the confined phase. The dashed black line is the result obtained for an extremal probe ( $\kappa = 0$ ). On the right there is a plot for the same values of  $\kappa$  in the deconfined phase, previously obtained in [26].

We also analyse the free energy difference  $\Delta\mathcal{F}$  in both phases. This is depicted in the figures below, in which we have represented with a dashed line the result obtained for  $\kappa = 0$  which can be found analytically,

$$\Delta\mathcal{F}|_{\kappa=0} = \sqrt{\lambda} kT \left( \gamma - \frac{\sqrt{2} \pi^{3/2} (1 - \gamma \hat{\sigma}_0^4)}{\hat{\sigma}_0 \Gamma(\frac{1}{4})^2} {}_2F_1 \left( \frac{1}{2}, \frac{3}{4}, \frac{1}{4}; \gamma \hat{\sigma}_0^4 \right) \right) . \quad (2.48)$$

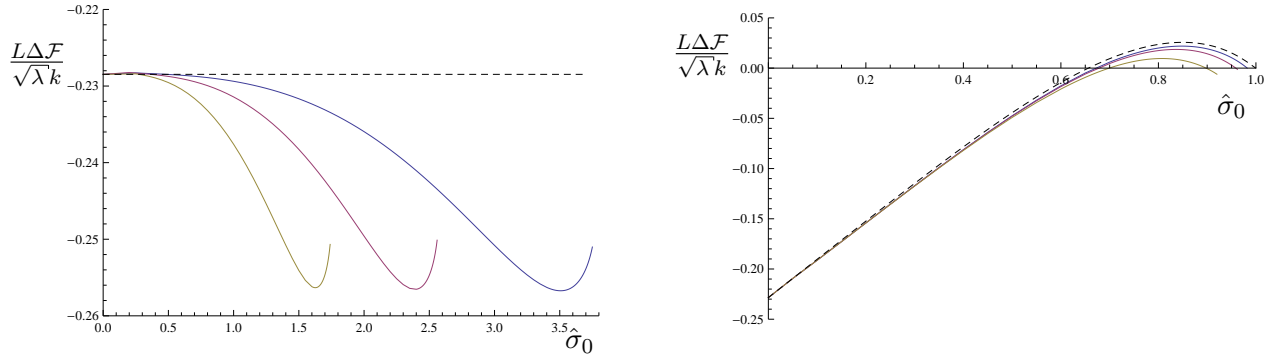


Figure 3: On the left is the plot of the free energy difference  $\frac{L\Delta\mathcal{F}}{\sqrt{\lambda}k}$  as a function of  $\hat{\sigma}_0$  obtained numerically. The dashed black line is the result obtained for an extremal probe ( $\kappa = 0$ ). On the right there is a plot for the same values of  $\kappa$  in the deconfined phase. The colour coding is the same as for the distances  $LT$  presented in the previous figures.

We can see that black strings in thermal AdS do not exhibit any phase transition to the Polyakov loop and hence are always the preferred configuration at a given temperature  $T$ . On the other hand, in the deconfined phase, a transition occurs around the value  $\hat{\sigma}_0 \sim 0.65$ . This describes the onset of the Debye screening effect.<sup>11</sup>

<sup>11</sup>See Ref. [26] for a detailed description of the deconfined phase.

In the thermal AdS background we see that the differences between the extremal probe method and the black probe method are significant as the free energy varies significantly compared to the constant value obtained for  $\kappa = 0$ . One should note that in thermal AdS, as in the case of a black hole background, there is a double scaling limit ( $\kappa \rightarrow 0$ ) that makes the probe extremal while still being in thermodynamical equilibrium with the finite temperature background. In thermal AdS, however, this limit leads to the same results as if it were probing AdS at zero-temperature, hence highlighting the qualitative new features between the extremal and black probe methods.

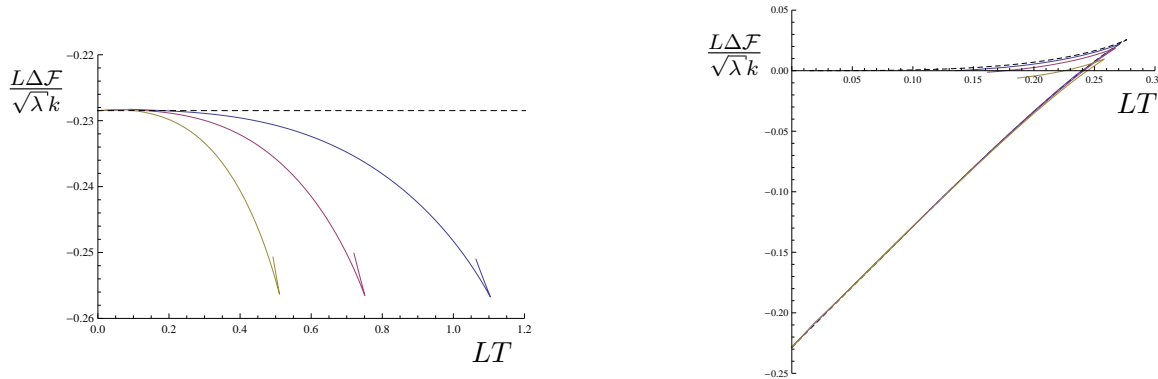


Figure 4: On the left is the plot of the free energy difference  $\frac{L\Delta\mathcal{F}}{\sqrt{\lambda}k}$  as a function of  $LT$  obtained numerically. The dashed black line is the result obtained for an extremal probe ( $\kappa = 0$ ). On the right there is a plot for the same values of  $\kappa$  in the deconfined phase.

One can also plot the free energy differences with respect to  $LT$ . This is depicted in the figures above. This again confirms that there is no phase transition occurring in thermal AdS while in the black hole background one finds the Debye screening length [26]

$$(LT)|_{\gamma=1, \Delta\mathcal{F}=0} \sim 0.240038 + 0.0379706\sqrt{\kappa} \quad . \quad (2.49)$$

The analysis presented here will serve as a means of comparison with our black string constructions in thermal Schrödinger in the following section.

### 3 Black strings in Schrödinger

In this section we construct three different types of black string configurations in thermal Schrödinger, two with spatially separated boundary endpoints and another with null separated boundary endpoints. The first configuration interpolates between the black string in AdS constructed in the previous section and a second configuration, which has no AdS counterpart, as the deformation parameter  $\ell$  is continuously increased.

The background we consider has  $\text{Sch}_5 \times S^5$  asymptotics and is obtained via a null Melvin twist of the extremal D3-brane and subsequently taking the near-horizon limit [3, 4, 5]. The resulting spacetime is



Schrödinger with dynamical exponent two<sup>12</sup> and metric<sup>13</sup>

$$ds^2 = \frac{R^2}{z^2} \left[ -\frac{\ell^2}{z^2} dt^2 + 2dt d\xi + dz^2 + dx_i^2 \right] + R^2 d\Omega_{(5)}^2 \quad , \quad (3.1)$$

where  $\xi$  is the boundary null coordinate and  $i = 1, 2$ . The boundary is located at  $z = 0$  and the origin at  $z = \infty$ . The parameter  $\ell$  is defined as  $\ell = \beta R^2$  with  $\beta$  having dimensions of inverse length. The choice  $\beta = 0$  brings us to AdS in null coordinates. Either by a boost in the  $(t, \xi)$  plane or by a uniform rescaling of all the  $\text{Sch}_5$  coordinates (these two operations differ by the characteristic anisotropic dilatation symmetry of  $\text{Sch}_5$ ), the parameter  $\ell$  can be scaled to any particular value. When placing a configuration at finite temperature  $T$  or with characteristic length  $L$  in (3.1), this implies that physical quantities are characterised by the dimensionless temperature  $\mathbf{T}$  and the dimensionless length  $\mathbf{L}$  defined as

$$\mathbf{T} \equiv T\ell \quad , \quad \mathbf{L} \equiv \frac{L}{\ell} \quad . \quad (3.2)$$

In this sense, the parameter  $\ell$  introduces a new reference scale into the problem, as the black hole horizon introduced a scale via its mass in AdS. The null Melvin twist also generates a  $B_{[2]}$  field of the form

$$B_{[2]} = 2\ell \frac{R^2}{z^2} (d\chi + \mathcal{A}) \wedge dt \quad , \quad d\Omega_{(5)}^2 = ds_{\mathbb{P}^2}^2 + (d\chi + \mathcal{A})^2 \quad , \quad (3.3)$$

where  $\mathcal{A}$  is the 1-form potential for the Kahler form on  $\mathbb{P}^2$ . However, we will always consider configurations at fixed coordinates in the five-sphere and hence couplings to the field  $B_{[2]}$  are absent throughout this paper.

We also note that we will only consider “thermal Schrödinger”, the Schrödinger analogue of “thermal AdS”, and not consider the black hole background with  $\text{Sch}_5 \times S^5$  asymptotics that one obtains via the null Melvin twist of the non-extremal black D3-brane [3, 4, 5] because in this case a non-trivial dilaton is generated in the process and how exactly a black probe couples to a background dilaton field is still to be worked out.<sup>14</sup> In [13] it was shown that in  $\text{Sch}_5 \times S^5$  there is a Schrödinger analogue of the Hawking-Page transition to a black hole background (from global Schrödinger [11] to the null Melvin twist (TsT dual) of the (asymptotically plane wave AdS) MMT black hole [3]). As long as we are at sufficiently small temperatures, the preferred phase will be thermal Schrödinger. We will then compare our results with those for thermal AdS black strings found in the previous section.

### 3.1 Interpolating black strings with spacelike separated boundary endpoints

In this section we construct the Schrödinger geometry for which its AdS counterpart was analysed in the previous section. This geometry interpolates between the AdS result, where  $\ell = 0$ , and a purely Schrödinger geometry, which will be studied in detail in the next section, when  $\ell$  is very large and the spacetime is highly deformed compared to AdS. In order to do so we introduce a new set of coordinates  $t', \xi'$  such that

$$t' = \frac{1}{\sqrt{2}}(t - \xi) \quad , \quad \xi' = \frac{1}{\sqrt{2}}(t + \xi) \quad . \quad (3.4)$$

<sup>12</sup>Unfortunately, dynamical exponents for Schrödinger spacetimes are commonly denoted in the literature by  $z$ , so in this case we have  $z = 2$ .

<sup>13</sup>Here we have followed the conventions of [4], however we have taken the parameter  $\beta$  defined in [4] and rescaled it such that  $\beta \rightarrow 1/\sqrt{2}\beta$  to avoid keeping track of factors of 2 in the metric.

<sup>14</sup>These couplings will be dealt with in [38].

The metric (3.1) then takes the form

$$ds^2 = \frac{R^2}{z^2} \left[ -g(z)dt'^2 - \frac{\ell^2}{z^2} dt' d\xi' + \left( 1 - \frac{\ell^2}{2z^2} \right) d\xi'^2 + dz^2 + dx_i^2 \right] + R^2 d\Omega_{(5)}^2 \quad , \quad g(z) = 1 + \frac{\ell^2}{2z^2} \quad . \quad (3.5)$$

We see that if we take  $\ell \rightarrow 0$  we obtain the metric (2.14) with  $\gamma = 0$  where  $t'$  is identified with  $t$  and  $\xi'$  with  $x_3$ . We now want to consider the analogous configuration as in the previous section where the string is stretched along the  $z$ -direction from the boundary to the bulk at  $z = \sigma_0$  and back to the boundary at  $z = 0$ . For this we chose the embedding map

$$t' = \tau \quad , \quad z = \sigma \quad , \quad x_1 = x(\sigma) \quad , \quad \xi' = 0 \quad , \quad x_i = 0 \quad , \forall i = 2, 3 \quad , \quad d\Omega_{(5)} = 0 \quad , \quad (3.6)$$

which leads to the induced metric

$$\gamma_{ab} d\sigma^a d\sigma^b = \frac{R^2}{\sigma^2} (-g(\sigma) d\tau^2 + (1 + x'(\sigma)^2) d\sigma^2) \quad , \quad (3.7)$$

and reduces to (2.16) when  $\ell = 0$ . Furthermore the pullback of the timelike Killing vector field  $\partial_{t'}$  onto the world volume  $\mathcal{W}_2$  yields  $\mathbf{k} = (R/\sigma)\sqrt{g(\sigma)}$ . Indeed we see that if  $\ell = 0$  we get  $\mathbf{k} = R/\sigma$  which is the result obtained in AdS when  $\gamma = 0$ , while if we take  $\ell$  to be very large we get  $\mathbf{k} \sim R\ell/(\sigma^2)$  which gives rise to a qualitatively different behaviour. This latter case deserves special attention and will be analysed in detail in the next section.

Using the above induced metric in (2.12) we find the free energy

$$\mathcal{F}[x(\sigma)] = -2A \left( \frac{3}{2\pi T} \right)^6 \int_0^{\sigma_0} d\sigma \sqrt{1 + x'(\sigma)^2} Z(\sigma) \quad , \quad Z(\sigma) = \frac{R^8}{\sigma^8} g(\sigma)^{\frac{7}{2}} \frac{1 + 6 \sinh^2 \alpha}{\cosh^6 \alpha} \quad , \quad (3.8)$$

which when varied and imposing the conditions  $x'(\sigma) \rightarrow \infty$  as  $\sigma \rightarrow \sigma_0$  leads to the solution for  $x'(\sigma)$

$$x'(\sigma) = \left( \frac{Z(\sigma)^2}{Z(\sigma_0)^2} - 1 \right)^{-\frac{1}{2}} \quad . \quad (3.9)$$

These solutions are valid solutions when certain requirements, which will be derived in the next section, similar to those obtained in Sec. 2.2 are imposed. The case at  $T = 0$  already presents itself as a very interesting one and difficult to track analytically. In that case we have

$$A \left( \frac{3}{2\pi T} \right)^6 Z(\sigma)|_{T \rightarrow 0} = Q_1 \frac{R^2}{\sigma^2} \sqrt{g(\sigma)} \quad . \quad (3.10)$$

One may now attempt to obtain analytic expressions for the distance between the boundary endpoints in terms of the bulk depth. However, this is not possible to obtain analytically for all values of the deformation parameter but only in the case  $\ell^2 = 0$  and in the limit  $\ell^2 \rightarrow \infty$ . One may also attempt to expand the sought after quantities for small and large  $\ell^2$  but it turns out that such series expansions in  $\ell^2$  do not converge for the functions that we are dealing with.<sup>15</sup> Therefore, we need to analyse the problem numerically. For this we consider the regularised free energy, obtained by subtracting to (3.8) a piece

$$\mathcal{F}_{\text{sub}} = \sqrt{\lambda} \frac{k}{\pi} \int_0^{\sigma_c} \frac{d\sigma}{\sigma^2} \sqrt{g(\sigma)} (1 - X) \quad , \quad (3.11)$$

---

<sup>15</sup>The reason for this is that the spacetimes with  $\ell^2 > 0$  and  $\ell^2 < 0$  are very different.

where  $X$  was defined in (2.33) and  $X \rightarrow 0$  as  $T \rightarrow 0$ . The resulting free energy takes the form

$$\mathcal{F}_{\text{loop}} = \sqrt{\lambda} \frac{k}{\pi} \left( h(\sigma_0) + \int_0^{\sigma_0} \frac{d\sigma}{\sigma^2} \sqrt{g(\sigma)} \left( (1-X) \sqrt{1-x'(\sigma)^2} - 1 \right) \right), \quad (3.12)$$

where we have defined the function  $h(\sigma_0)$  as

$$h(\sigma_0) = -\frac{1}{2} \left( \frac{\sqrt{\ell^2 + 2\sigma_0^2}}{\sqrt{2}\sigma_0^2} + \sqrt{2} \frac{\text{arcsinh}(\frac{\ell}{\sqrt{2}\sigma_0})}{\ell} \right). \quad (3.13)$$

The Polyakov loop configuration consisting of two straight strings with  $x'(\sigma) = 0$ , following the same steps as in Sec. 2.2, can be argued to have zero free energy for any value of  $\ell, \kappa, T$ . We now obtain expressions for the length  $L$ , computed in a similar manner as in (2.41), and  $\mathcal{F}_{\text{loop}}$  as a function of  $\sigma_0$  at  $T = 0$ . This behaviour is depicted in the figures below.

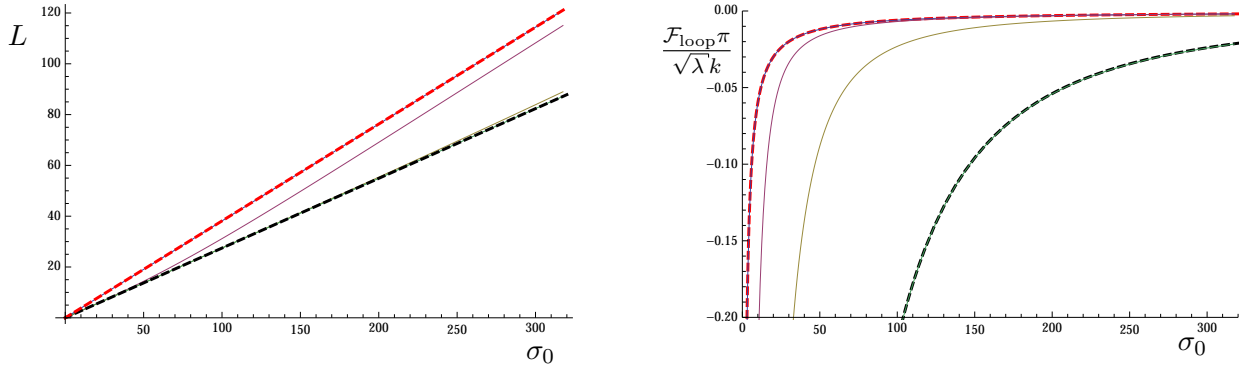


Figure 5: On the left we have the behaviour of the distance  $L$  as function of  $\sigma_0$  at  $T = 0$ . The dashed solid red line corresponds to the AdS result given by the first term in (2.41) while the dashed solid black line corresponds to the result (3.14). The blue line represents  $(\ell/\sqrt{2}) = 0.0001$ , the magenta line  $(\ell/\sqrt{2}) = 1$ , the yellow line  $(\ell/\sqrt{2}) = 1000$  and the green line  $(\ell/\sqrt{2}) = 10000$ . On the right we have the behaviour of the free energy as a function of  $\sigma_0$  at  $T = 0$ . The dashed solid red line is the first term obtained in (2.42) while the dashed solid black line is the result obtained in (3.14).

On the l.h.s. of the figure above we have shown the behaviour of the length  $L$  as a function of  $\sigma_0$  and on the r.h.s. the behaviour of the free energy as a function of  $\sigma_0$ . The dashed red line on the l.h.s. corresponds to the zero-temperature AdS result for the length  $L$  given in (2.41) while the dashed line on the r.h.s. corresponds to the leading order result for AdS given in (2.42). The dashed black lines depict the result obtained analytically in the limit  $\ell \rightarrow \infty$

$$L|_{\ell \rightarrow \infty} = 3\sqrt{\pi}\sigma_0 \frac{\Gamma(\frac{5}{3})}{\Gamma(\frac{1}{6})}, \quad \frac{\mathcal{F}_{\text{loop}}\pi}{\sqrt{\lambda}k}|_{\ell \rightarrow \infty} = \frac{\sqrt{\pi}\ell}{6\sigma_0^2} \frac{\Gamma(-\frac{1}{6})}{\Gamma(\frac{1}{6})} \quad (3.14)$$

which will be derived in the next section. We can see from the plots above that the behaviour for any value of  $\ell$  is bounded by the cases  $\ell = 0$  and  $\ell \rightarrow \infty$  and hence between an AdS geometry and a purely Schrödinger geometry. This means, as we will see explicitly below, that since  $\mathcal{F}_{\text{loop}}$  scales with  $\sigma_0^{-1}$  for  $\ell = 0$  and with  $\sigma_0^{-2}$  for very large  $\ell$  and since the behaviour of the curve  $(L, \sigma_0)$  is approximately linear (closer scrutiny reveals that for small values of  $\sigma_0$  there is a departure from the linear behaviour exhibited in Fig. 5), then the free energy is scaling like  $L^{-1}$  for  $\ell = 0$  and like  $L^{-2}$  for  $\ell \rightarrow \infty$ . Therefore we see that

in terms of the dual theory, the expectation value of these operators is changing substantially for large values of the deformation parameter.

We now consider the finite temperature case and for non-zero  $\kappa$ . For these purposes we introduce the dimensionless coordinate  $\hat{\sigma} = \pi T \sigma$  and define the parameter  $\kappa$  as in (2.23) but now with  $f(\hat{\sigma})$  replaced by  $g(\hat{\sigma})$  on the r.h.s. of (2.23). We further introduce the parameter  $\hat{\ell}$  such that

$$g(\hat{\sigma}) = 1 + \frac{\hat{\ell}^2}{\hat{\sigma}^2} \quad , \quad \hat{\ell}^2 = \frac{\pi^2 T^2 \ell^2}{2} \quad . \quad (3.15)$$

The qualitative behaviour of the dimensionless length  $LT$  and the free energy at finite temperature is qualitatively similar to that seen in Fig. 5, the difference being that the strings cannot extend all the way to the bulk but instead have a maximum distance as in (2.24) given by

$$\hat{\sigma}_c^2 = \frac{2^{\frac{1}{3}}}{5^{\frac{5}{6}}} \kappa^{-\frac{1}{3}} \left( 1 + \sqrt{1 + 2^{\frac{2}{3}} 5^{\frac{5}{6}} \hat{\ell}^2 \kappa^{\frac{1}{3}}} \right) \quad . \quad (3.16)$$

We now solve for the free energy numerically as a function of  $\hat{\sigma}_0$  for several values of  $\hat{\ell}$  and  $\kappa = 0.001$ . The result is depicted in the figure below. One can observe that, as in the case  $T = 0$  the results for any  $\hat{\ell}$  are bounded between the AdS result and the large  $\hat{\ell}$  result.

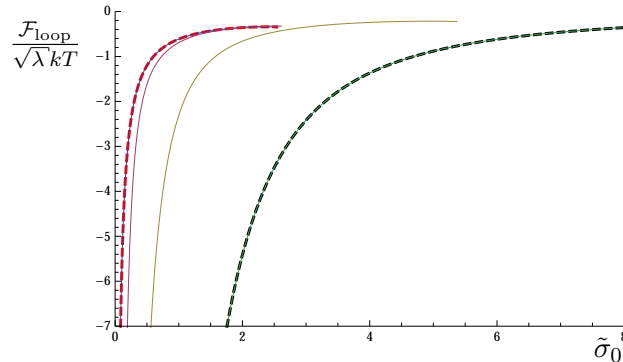


Figure 6: The free energy as a function of  $\hat{\sigma}_0$  for  $\kappa = 0.001$  obtained numerically. The red dashed line is the result obtained for  $\hat{\ell} = 0$  while the dashed black line is the result for  $\hat{\ell} = 100$  obtained using the results of the next section. The blue line corresponds to the case  $\hat{\ell} = 0.001$ , the magenta line to the case  $\hat{\ell} = 1$ , the yellow line to  $\hat{\ell} = 10$  and the green line to  $\hat{\ell} = 100$ .

The behaviour exhibited here is generic for any small value of  $\kappa$ . The behaviour of  $LT$  as a function of  $\hat{\sigma}_0$  is also approximately linear, except near  $\hat{\sigma}_c$ , and hence the dependence of  $\mathcal{F}_{\text{loop}}$  on  $LT$  shows approximately the same behaviour as what is depicted below. We will now analyse the case  $\hat{\ell} \rightarrow \infty$  in detail.

### 3.2 Black strings with spacelike separated boundary endpoints

In this section we construct another class of black strings with endpoints on a boundary spatial direction. This class of solutions, valid for any value of the deformation parameter  $\ell$ , meets the class of interpolating black strings studied in the previous section when  $\ell$  is taken to be very large. To that end, we go back to the coordinates  $(t, \xi)$ . As in the previous cases, we fix the endpoints of the string in the  $x_1$ -direction

and stretch it into the bulk along the  $z$ -direction in the background geometry (3.1). This leads to the embedding map

$$t = \tau \quad , \quad z = \sigma \quad , \quad x_1 = x(\sigma) \quad , \quad \xi = x_2 = d\Omega_{(5)} = 0 \quad , \quad (3.17)$$

and hence the problem is symmetric around  $x_1 = L/2$ . The induced metric on  $\mathcal{W}_2$ , using (3.1), is

$$\gamma_{ab}d\sigma^a d\sigma^b = \frac{R^2}{\sigma^2} \left( -\frac{\ell^2}{\sigma^2} d\tau^2 + (1 + x'(\sigma)^2) d\sigma^2 \right) \quad . \quad (3.18)$$

The pull-back of the background time-like Killing vector field  $\partial_t$  leads to a world volume timelike Killing vector field  $\partial_\tau$  with norm  $R_0 = \mathbf{k} = R\ell/\sigma^2$ . This different power of  $\sigma$  when compared with the case of Sec. 2.2 will lead to different local properties of the black string. In particular we see that when  $\ell \rightarrow 0$  we get  $\mathbf{k} \rightarrow 0$  and hence these solutions have no AdS counterpart. In terms of the dual DLCQ $_\beta$  field theory, these presumably correspond to operators that are well defined in the  $\beta$ -deformed DLCQ theory but are singular in the conventional DLCQ of the theory (due to the notorious zero mode problem, see e.g. [40, 41]).

The free energy takes the same form as in (2.17) but with a different function  $H(\sigma)$

$$\mathcal{F}[x(\sigma)] = -2A \left( \frac{3}{2\pi T} \right)^6 \int_0^{\sigma_0} d\sigma \sqrt{1 + x'(\sigma)^2} H(\sigma) \quad , \quad H(\sigma) = \frac{R^8 \ell^7}{\sigma^{15}} \frac{1 + 6 \sinh^2 \alpha}{\cosh^6 \alpha} \quad , \quad (3.19)$$

where the overall factor of 2 in the free energy is again due to the fact that the string is stretched all the way from  $(x_1, z) = (0, 0)$  to  $(x_1, z) = (L/2, \sigma_0)$  and back again to  $(x_1, z) = (L, 0)$ . The equation of motion that follows from varying (3.19) is

$$\left( \frac{x'(\sigma)}{\sqrt{1 + x'(\sigma)^2} H(\sigma)} \right)' = 0 \quad . \quad (3.20)$$

As in the AdS case, we consider solutions for which  $x'(\sigma) > 0$  and which satisfy the boundary conditions  $x(0) = 0$  and  $x'(\sigma) \rightarrow \infty$  for  $\sigma \rightarrow \sigma_0$ . This gives rise to the solution

$$x'(\sigma) = \left( \frac{H(\sigma)^2}{H(\sigma_0)^2} - 1 \right)^{-\frac{1}{2}} \quad . \quad (3.21)$$

We note that in thermal Schrodinger there is also an analogous solution to the Polyakov loop in AdS, characterised by two straight strings stretched from the boundary to the bulk with  $x'(\sigma) = 0$ , trivially satisfying (3.20). The boundary conditions (2.2) must also be satisfied. In this case, the unit normalised, orthogonal covector to the world volume boundary is

$$\eta_b d\sigma^b = \frac{R}{\sigma} \sqrt{1 + x'(\sigma)^2} d\sigma \quad . \quad (3.22)$$

Evaluating (2.2) explicitly at the boundary  $\sigma = 0$  leads to

$$T^{\sigma\sigma} \eta_\sigma|_{\sigma=0} = -Q_1 \frac{(1 + 6 \sinh^2 \alpha)}{6 \sinh \alpha \cosh \alpha} \frac{\sigma}{R} = 0 \quad , \quad J^{\tau\sigma} \eta_\sigma|_{\sigma=0} = Q_1 \frac{\sigma^2}{R\ell} = 0 \quad , \quad (3.23)$$

where we have used the fact that  $H(\sigma) \rightarrow \infty$  when  $\sigma \rightarrow 0$ . The same result holds for the straight string.

### Critical distance and regime of validity

As in (2.23) one can define the parameter  $\kappa$  via the relation

$$\kappa = \frac{2^5 Q_1}{3^7 A R^6} = \frac{1}{\tilde{\sigma}^{12}} \frac{\sinh \alpha}{\cosh^5 \alpha} \quad , \quad \tilde{\sigma} = \sqrt{\pi \frac{T}{\ell}} \sigma \quad . \quad (3.24)$$

The rescaling of the coordinate  $\sigma$  was chosen such that  $\kappa$  has no dependence on the temperature and remains the same as in the AdS case. This leads us to the result for the critical distance  $\tilde{\sigma}_c$

$$\tilde{\sigma}_c^2 = \frac{2^{\frac{2}{3}}}{5^{\frac{5}{12}}} \kappa^{-\frac{1}{6}} \quad . \quad (3.25)$$

The largest scale characterising the probe is  $r_c = R\kappa^{1/6}$  and the validity of the supergravity approximation still requires (2.28), while the requirement  $r_c \ll R$  again yields  $\kappa \ll 1$ . Furthermore, one must have that the local temperature variations compared to the background temperature must be small, i.e.,

$$r_c \frac{T'}{T} = 2r_c \frac{\sigma}{R\ell} \ll 1 \quad . \quad (3.26)$$

Near the boundary this is trivially satisfied, while for  $\sigma \sim \sigma_c$  we obtain the requirement

$$\kappa \ll \left( \frac{\ell}{T} \right)^6 \quad , \quad (3.27)$$

which just implies a very mild bound on the ratio  $\ell/T$ , in particular, for low temperatures this is easily satisfied. One further needs to require that the string is thin compared to the extrinsic curvature length scale  $L_{\text{ext}}$  given by

$$L_{\text{ext}}(\sigma) = |K^\rho N_\rho|^{-1} = \frac{R}{\sigma} \frac{(1 + x'(\sigma)^2)^{3/2}}{|x''(\sigma)|} \quad . \quad (3.28)$$

The minimum curvature is attained when  $\sigma = \sigma_0$  and requiring  $r_c \ll L_{\text{ext}}(\sigma_0 = \sigma_c)$  just results in  $r_c \ll R$  which has already been imposed. We thus conclude that these black strings in thermal Schrödinger are valid solutions all the way up to  $\sigma = \sigma_c$ . We further note that the Polyakov loop in the case  $\kappa \neq 0$ , as in AdS, lies outside the regime of validity of the method we are employing.

### Regularized free energy and large $N$ expansion

The free energy (3.19) using (2.7) and the AdS/CFT dictionary can be written as

$$\mathcal{F}[x(\tilde{\sigma})] = \sqrt{\lambda} k T \int_0^{\tilde{\sigma}_0} \frac{d\tilde{\sigma}}{\tilde{\sigma}^3} (1 - X) \sqrt{1 + x'(\tilde{\sigma})^2} \quad , \quad (3.29)$$

where  $X$  was defined in (2.33), and diverges at the boundary  $\sigma = 0$ . Applying the same prescription as for AdS, we subtract a piece

$$\mathcal{F}_{\text{sub}} = \sqrt{\lambda} k T \int_0^{\tilde{\sigma}_{\text{cut}}} \frac{d\tilde{\sigma}}{\tilde{\sigma}^3} (1 - X) \quad , \quad (3.30)$$

such that the free energy difference  $\Delta\mathcal{F}$  can be written as  $\Delta\mathcal{F} = \mathcal{F} - \mathcal{F}_{\text{sub}} = \mathcal{F}_{\text{loop}} - 2\mathcal{F}_{\text{P}}$ , where

$$\mathcal{F}_{\text{loop}} = \sqrt{\lambda} k T \left( -\frac{1}{2\tilde{\sigma}_0^2} + \int_0^{\tilde{\sigma}_0} \frac{d\tilde{\sigma}}{\tilde{\sigma}^3} \left( (1 - X) \sqrt{1 + x'(\tilde{\sigma})^2} - 1 \right) \right) \quad , \quad (3.31)$$

$$\mathcal{F}_P = -\frac{1}{2}\sqrt{\lambda}kT \left( \frac{1}{2\tilde{\sigma}_{\text{cut}}^2} + \int_0^{\tilde{\sigma}_{\text{cut}}} \frac{d\tilde{\sigma}}{\tilde{\sigma}^3} X \right) . \quad (3.32)$$

Using similar arguments as in the AdS case, relying on the independence of  $\kappa$  on  $T$  and demanding the results at  $T = 0$  to coincide for both the extremal and black probes, one concludes that  $\mathcal{F}_P = 0$  for any  $\kappa$ . In order to obtain analytic results, we now proceed and make a similar expansion as in (2.39) and solve for  $\phi$  obtaining

$$\cosh^2 \alpha = \frac{1}{\tilde{\sigma}^6 \sqrt{\kappa}} - \frac{1}{4} - \frac{5}{32} \tilde{\sigma}^6 \sqrt{\kappa} + \mathcal{O}(\kappa) . \quad (3.33)$$

With this in hand we can compute the total distance between the boundary endpoints to order  $\mathcal{O}(\kappa)$

$$\mathbf{L}\mathbf{T}^{\frac{1}{2}} = \frac{2}{\sqrt{\pi}} \int_0^{\tilde{\sigma}_0} \frac{dx(\tilde{\sigma})}{d\tilde{\sigma}} = 3\tilde{\sigma}_0 \frac{\Gamma(\frac{5}{3})}{\Gamma(\frac{1}{6})} (1 - \tilde{\sigma}_0^6 \sqrt{\kappa} + \mathcal{O}(\kappa)) , \quad (3.34)$$

where the dimensionless parameters  $\mathbf{L}$  and  $\mathbf{T}$  were introduced in Eq. (3.2). The leading order result in (3.34), when written in terms of  $\sigma_0$  leads to the result presented on the l.h.s. of (3.14). Similarly, we can expand (3.31) in powers of  $\kappa$ , obtaining

$$\mathcal{F}_{\text{loop}} = \sqrt{\lambda}kT \frac{\sqrt{\pi}}{6\tilde{\sigma}_0^2} \frac{\Gamma(-\frac{1}{3})}{\Gamma(\frac{1}{6})} \left( 1 + \frac{4}{3} \tilde{\sigma}_0^6 \sqrt{\kappa} + \mathcal{O}(\kappa) \right) , \quad (3.35)$$

for which the leading order term, when written in terms of  $\sigma_0$  yields the result presented on the r.h.s. of (3.14). Inverting now (3.34), introducing in (3.35) and expanding for small  $\kappa$  and small  $L$  we find

$$\mathcal{F}_{\text{loop}} = -2\sqrt{\pi} \frac{\sqrt{\lambda}k}{\mathbf{L}^2} \left( \frac{\Gamma(-\frac{4}{3})\Gamma(\frac{5}{3})^2}{\Gamma(\frac{1}{6})^3} - \frac{\Gamma(\frac{1}{6})^3}{96\Gamma(\frac{2}{3})^3} \mathbf{L}^6 \mathbf{T}^3 + \mathcal{O}(\kappa) + \mathcal{O}(\mathbf{L}^{10}) \right) , \quad (3.36)$$

where we have defined  $\mathcal{F}_{\text{loop}} = \mathcal{F}_{\text{loop}} \ell$ . We see that the free energy has a different structure when compared with the free energy in AdS (2.43). In particular it scales with  $\mathbf{L}^{-2}$ , as advertised in the previous section, while temperature corrections come in powers of  $\mathbf{L}\mathbf{T}^{\frac{1}{2}}$ . The combination  $\mathcal{F}_{\text{loop}} \mathbf{L}^2$  is scale invariant and also  $\mathbf{L}\mathbf{T}^{\frac{1}{2}}$ , therefore we see that the quark-antiquark potential in  $\text{Sch}_5 \times S^5$  is capturing the conformal invariance of the dual non-relativistic CFT.

With (3.36) in hand, we can evaluate the black string entropy and mass using (2.13). Specifically, we find

$$S = 2\sqrt{\pi} \frac{\sqrt{\lambda}k}{\mathbf{T}\mathbf{L}^2} \left( -\frac{\Gamma(\frac{1}{6})^3}{16\Gamma(\frac{2}{3})^3} \mathbf{L}^6 \mathbf{T}^3 + \mathcal{O}(\kappa) + \mathcal{O}(\mathbf{L}^{10}) \right) . \quad (3.37)$$

$$\mathbf{M} = -2\sqrt{\pi} \frac{\sqrt{\lambda}k}{\mathbf{L}^2} \left( \frac{\Gamma(-\frac{4}{3})\Gamma(\frac{5}{3})^2}{\Gamma(\frac{1}{6})^3} - \frac{7\Gamma(\frac{1}{6})^3}{96\Gamma(\frac{2}{3})^3} \mathbf{L}^6 \mathbf{T}^3 + \mathcal{O}(\kappa) + \mathcal{O}(\mathbf{L}^{10}) \right) \quad (3.38)$$

where we have defined  $\mathbf{M} = M\ell$ .

We now use a mixture of analytic and numerical methods to obtain the solution space of these black strings. We begin by solving for the distance  $\mathbf{L}\mathbf{T}^{\frac{1}{2}}$  numerically, which is depicted on the l.h.s. of the figure below.

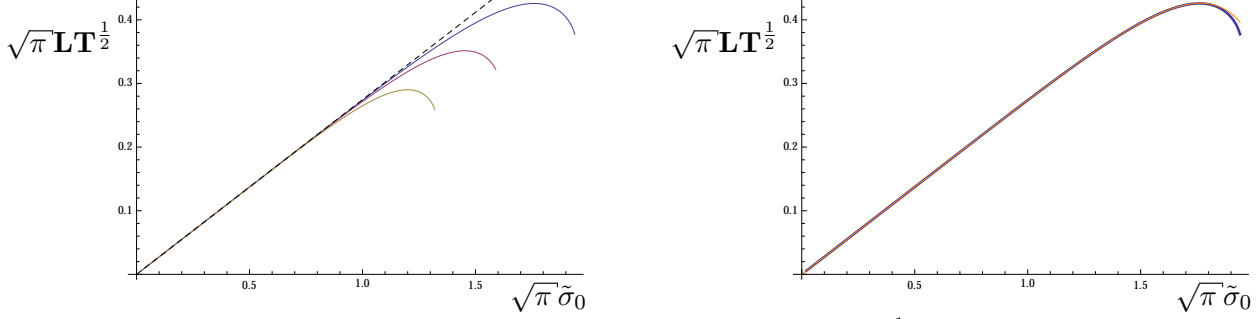


Figure 7: On the left is the plot of the dimensionless distance  $\sqrt{\pi} \mathbf{L} \mathbf{T}^{\frac{1}{2}}$  as a function of  $\sqrt{\pi} \tilde{\sigma}_0$  for  $\kappa = 0.01$  (yellow line),  $\kappa = 0.001$  (red line) and  $\kappa = 0.0001$  (blue line). The dashed black line is the case  $\kappa = 0$ . On the right hand side we have  $\sqrt{\pi} \mathbf{L} \mathbf{T}^{\frac{1}{2}}$  as a function of  $\sqrt{\pi} \tilde{\sigma}_0$ , where the thick blue line represents  $\kappa = 0.0001$  while thin orange line represents the analytic expansion to order  $\mathcal{O}(\kappa^{5/2})$  given in (3.39).

The dashed black line in the l.h.s. of the figure above represents the result for an extremal probe obtained by setting  $\kappa = 0$  in (3.34). The dependence of the dimensionless distance on the bulk depth  $\tilde{\sigma}_0$  is similar to that seen in thermal AdS in Fig. 2. In order to describe these curves analytically for small  $\kappa$  we can expand the dimensionless distance to  $\mathcal{O}(\kappa^{5/2})$  in order to find

$$\mathbf{L} \mathbf{T}^{\frac{1}{2}} = 3\tilde{\sigma}_0 \frac{\Gamma(\frac{5}{3})}{\Gamma(\frac{1}{6})} \left( 1 - \tilde{\sigma}_0^6 \sqrt{\kappa} - \frac{83}{504} \tilde{\sigma}_0^{12} \kappa - \frac{10471}{78624} \tilde{\sigma}_0^{18} \kappa^{3/2} - \frac{820955}{5975424} \tilde{\sigma}_0^{24} \kappa^2 + \mathcal{O}(\kappa^{5/2}) \right) , \quad (3.39)$$

where we have used the fact that to this order

$$\cosh^2 \alpha = \frac{1}{\tilde{\sigma}_0^6 \sqrt{\kappa}} - \frac{1}{4} - \frac{5}{32} \tilde{\sigma}_0^6 \sqrt{\kappa} - \frac{5}{32} \tilde{\sigma}_0^{12} \kappa - \frac{385}{2048} \tilde{\sigma}_0^{18} \kappa^{3/2} - \frac{19}{128} \tilde{\sigma}_0^{24} \kappa^2 + \mathcal{O}(\kappa^{5/2}) . \quad (3.40)$$

Using (3.39) we have drawn the thin orange curve in the r.h.s. of the figure above. As seen from the plot, the expansion (3.39) captures all the essentially features accurately up to values  $\sigma \sim \sigma_c$ . This behaviour is generic for small values of  $\kappa$ . We can perform a similar expansion in the free energy (3.31) obtaining

$$\mathcal{F}_{\text{loop}} = \sqrt{\lambda} k T \frac{\sqrt{\pi} \Gamma(-\frac{1}{3})}{6\tilde{\sigma}_0^2 \Gamma(\frac{1}{6})} \left( 1 + \frac{4}{3} \tilde{\sigma}_0^6 \sqrt{\kappa} + \frac{25}{84} \tilde{\sigma}_0^{12} \kappa + \frac{7955}{39312} \tilde{\sigma}_0^{18} \kappa - \frac{51805}{689472} \tilde{\sigma}_0^{24} \kappa^2 + \mathcal{O}(\kappa^{5/2}) \right) . \quad (3.41)$$

Below we show the behaviour of  $\mathcal{F}_{\text{loop}}$  using (3.41) and (3.39) as a function of the bulk depth and as a function of the dimensionless distance between boundary endpoints.

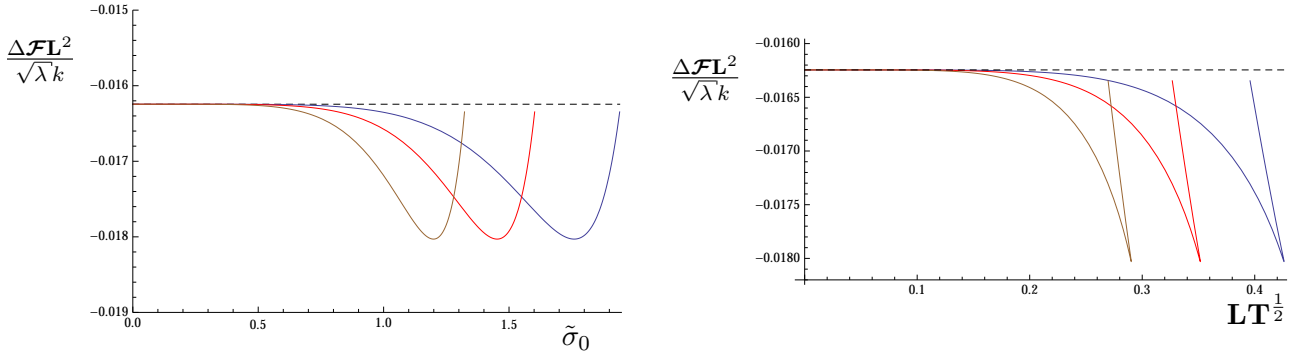


Figure 8: On the left is the plot of  $\frac{\Delta \mathcal{F} \mathbf{L}^2}{\sqrt{\lambda} k}$ , where  $\Delta \mathcal{F} \equiv \mathcal{F}_{\text{loop}} - \mathcal{F}_W \ell$ , as function of  $\tilde{\sigma}_0$  and as a function of  $\mathbf{L} \mathbf{T}^{\frac{1}{2}}$  on the right. The colour coding is the same as on the l.h.s. of the previous figure. The dashed line is the result for  $\kappa = 0$ .



The black dashed line depicted above is the result for  $\kappa = 0$  which using Eq. (3.41) is given by

$$\frac{\Delta \mathcal{F} \mathbf{L}^2}{\sqrt{\lambda} k} \big|_{\kappa=0} = \frac{3\sqrt{\pi}}{2\tilde{\sigma}_0^2} \frac{\Gamma(-\frac{1}{3})\Gamma(\frac{5}{3})^2}{\Gamma(\frac{1}{6})^3} \quad (3.42)$$

From the above figures, it is clear that, as in thermal AdS, there is no phase transition in thermal Schrödinger and hence the black string configuration is always preferred compared to the Polyakov loop. When working in thermal Schrödinger spacetimes we will use the term “phase transition” in this sense, i.e. we analyse possible phase transitions representing the onset of Debye screening effects, by which we mean transitions from Wilson loops to Polyakov loops.

### 3.3 Black strings with lightlike separated boundary endpoints

In this section we consider black strings with lightlike separated boundary endpoints which have no AdS counterpart. In this case we fix the endpoints of the string in the null  $\xi$ -direction of the metric (3.1) and stretch it into the bulk along the  $z$ -direction. The embedding map is now

$$t = \tau \quad , \quad z = \sigma \quad , \quad \xi = \xi(\sigma) \quad , \quad x_1 = x_2 = d\Omega_{(5)} = 0 \quad , \quad (3.43)$$

rendering the problem symmetric around  $\xi = L/2$ . With this we find the induced metric which is now stationary

$$\gamma_{ab} d\sigma^a d\sigma^b = \frac{R^2}{\sigma^2} \left( -\frac{\ell^2}{\sigma^2} d\tau^2 + 2\xi'(\sigma) d\tau d\sigma + d\sigma^2 \right) \quad , \quad (3.44)$$

and the corresponding free energy

$$\mathcal{F}[\xi(\sigma)] = -2A \left( \frac{3}{2\pi T} \right)^6 \int_0^{\sigma_0} d\sigma \sqrt{1 + \frac{\sigma^2}{\ell^2} \xi'(\sigma)^2} H(\sigma) \quad , \quad (3.45)$$

where  $H(\sigma)$  was given in (3.19). Here, we have used the fact that the norm of the Killing vector field  $\partial_t$  on the world volume is  $\mathbf{k} = R\ell/\sigma^2$ . The equations of motion that follow from (3.45) are

$$\left( \frac{\sigma^2 \xi'(\sigma)}{\ell^2 \sqrt{1 + \frac{\sigma^2}{\ell^2} \xi'(\sigma)^2}} H(\sigma) \right)' = 0 \quad . \quad (3.46)$$

We look for configurations that stretch from the boundary to  $\sigma_0$  and back again to the boundary and that satisfy  $\xi'(\sigma) > 0$  as well as  $\xi(\sigma) \rightarrow \infty$  as  $\sigma \rightarrow \sigma_0$ . In this case we obtain the solution

$$\xi'(\sigma) = \left( \frac{\sigma^4 H^2(\sigma)}{\ell^2 \sigma_0^2 H^2(\sigma_0)} - \frac{\sigma^2}{\ell^2} \right)^{-\frac{1}{2}} \quad . \quad (3.47)$$

We note that the equation of motion also admits a solution of a stretched string  $\xi'(\sigma) = 0$  analogous to the Polyakov loop.

The boundary conditions (2.2) are also satisfied. In order to see this explicitly, we construct the normalised orthogonal covector to the world volume boundary

$$\eta_b d\sigma^b = \frac{R}{\sigma} \sqrt{1 + \frac{\sigma^2}{\ell^2} \xi'(\sigma)^2} d\sigma \quad . \quad (3.48)$$

Evaluating the boundary conditions (2.2) explicitly leads to the same results as in (3.23), however in this case we have an extra component of the stress-energy tensor that must also be considered

$$T^{\tau\sigma}\eta_\sigma|_{\sigma=0} = -Q_1 \frac{(1 + 6 \sinh^2 \alpha)}{6 \sinh \alpha \cosh \alpha} \frac{\sigma^3}{R\ell^2} \xi'(\sigma) = 0 \quad , \quad (3.49)$$

where we have used the fact that  $\xi'(\sigma) \rightarrow 0$  when  $\sigma \rightarrow 0$ . The same result holds for the straight string.

We have established the existence of this type of black strings in  $\text{Sch}_5 \times S^5$  but we need to make sure that they lie within the regime of validity of our approximation. This analysis is in fact the same as in Sec. 3.2, in particular the critical distance  $\tilde{\sigma}_c$  remains unaltered. The only difference resides in the specific details of the extrinsic curvature length scale. In this case we find

$$L_{\text{ext}}(\sigma) = R \frac{(1 + \xi'(\sigma)^2)^{3/2}}{|\xi'(\sigma) + \xi'(\sigma)^3 - \sigma \xi''(\sigma)|} \quad . \quad (3.50)$$

The minimum curvature scale is found at  $\sigma = \sigma_0$ . It is easy to check that for small  $\kappa$  either for  $\sigma_0 \sim 0$  or  $\sigma_0 \sim \sigma_c$  we find that  $L_{\text{ext}}(\sigma_0) \sim R$ . This requires that  $\kappa \ll 1$ , which has already been imposed. We conclude that these black strings are valid solutions all the way up to  $\sigma_c$ .

### Regularized free energy and large $N$ expansion

As in the previous cases, the free energy (3.45) can be rewritten as

$$\mathcal{F}[\xi(\tilde{\sigma})] = \sqrt{\lambda} kT \int_0^{\tilde{\sigma}_0} \frac{d\tilde{\sigma}}{\tilde{\sigma}^3} (1 - X) \sqrt{1 + \tilde{\sigma}^2 \xi'(\tilde{\sigma})^2} \quad , \quad (3.51)$$

and diverges at  $\tilde{\sigma} = 0$ . In order to deal with this divergence we subtract the piece (3.30) obtaining the regularised free energy

$$\mathcal{F}_{\text{loop}} = \sqrt{\lambda} kT \left( -\frac{1}{2\tilde{\sigma}_0^2} + \int_0^{\tilde{\sigma}_0} \frac{d\tilde{\sigma}}{\tilde{\sigma}^3} \left( (1 - X) \sqrt{1 + \tilde{\sigma}^2 \xi'(\tilde{\sigma})^2} - 1 \right) \right) \quad , \quad (3.52)$$

while  $\mathcal{F}_W$  in (3.32) still vanishes using the same type of analysis. We can now use the expansion (3.33) in order to evaluate the distance between the string endpoints

$$\mathbf{L} = 2 \int_0^{\tilde{\sigma}_0} d\tilde{\sigma} \frac{\partial \xi(\tilde{\sigma})}{\partial \tilde{\sigma}} = \frac{\pi}{2} - \frac{2\tilde{\sigma}_0^6}{3} \sqrt{\kappa} + \mathcal{O}(\kappa) \quad . \quad (3.53)$$

This expression is significantly different than what we have encountered before. First of all, in the case of an extremal probe ( $\kappa = 0$ ) the distance between the endpoints does not depend on how much the string is stretched into the bulk. Secondly, the distance decreases for increasing bulk depth  $\tilde{\sigma}_0$  and non-zero  $\kappa$ . Thirdly, it does not depend on the temperature  $T$ . Performing a similar expansion in (3.52), inverting (3.53) and introducing in (3.52) leads to

$$\mathcal{F}_{\text{loop}} = -\sqrt{\lambda} kT \left( \frac{9(\pi - 2\mathbf{L})^2}{2^7} \right)^{\frac{1}{3}} \kappa^{\frac{1}{6}} + \mathcal{O}(\kappa) \quad , \quad (3.54)$$

which is again qualitatively different than in the previous cases, namely, it is temperature dependent and the first correction is proportional to  $\kappa^{\frac{1}{6}}$  and vanishes for zero  $\kappa$ . This configuration is thus more sensitive

to corrections due to the internal degrees of freedom of the probe. This indicates that black probes can be used to distinguish Schrödinger spacetimes from AdS in null coordinates.<sup>16</sup>

We can also evaluate the black string entropy and mass using (3.54) and (2.13),

$$S = \sqrt{\lambda} k \left( \frac{9(\pi - 2\mathbf{L})^2}{2^7} \right)^{\frac{1}{3}} \kappa^{\frac{1}{6}} + \mathcal{O}(\kappa) \quad , \quad \mathbf{M} = -\sqrt{\lambda} k \mathbf{T} \left( \frac{9(\pi - 2\mathbf{L})^2}{2^6} \right)^{\frac{1}{3}} \kappa^{\frac{1}{6}} + \mathcal{O}(\kappa) \quad , \quad (3.55)$$

yielding an entropy which is finite, independent of  $T$  and vanishing when  $\kappa = 0$ .

In order to depict the solution space, we solve numerically for the length  $\mathbf{L}$  obtaining the figure below, which is accurately reproduced up to values  $\sigma \sim \sigma_c$  and for small  $\kappa$  by the following expansion

$$\begin{aligned} \mathbf{L} = & \frac{\pi}{2} - \frac{2}{3} \tilde{\sigma}_0^6 \sqrt{\kappa} + \frac{1}{384} (-256 + 55\pi) \tilde{\sigma}_0^{12} \kappa - \frac{5(7856 - 2079\pi)}{36288} \tilde{\sigma}_0^{18} \kappa^{3/2} \\ & + \frac{5(2063061\pi - 7178240) \tilde{\sigma}_0^{24}}{9289728\pi} \kappa^2 + \mathcal{O}(\kappa^{5/2}) \quad . \end{aligned} \quad (3.56)$$

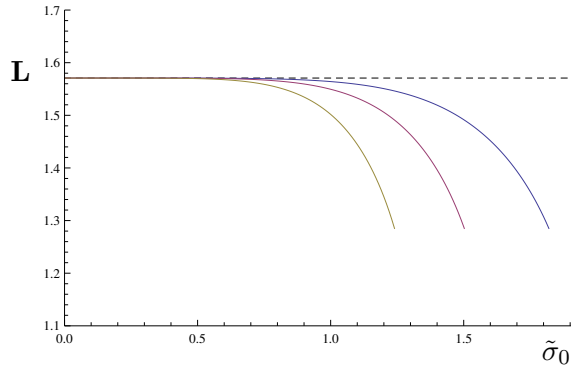


Figure 9: The dimensionless distance  $\mathbf{L}$  numerically obtained as a function of  $\tilde{\sigma}_0$  for  $\kappa = 0.01$  (yellow line),  $\kappa = 0.001$  (red line) and  $\kappa = 0.00001$  (blue line). The dashed line represents the result for  $\kappa = 0$ .

As we can see from the figure above, the maximum length is achieved when  $\kappa = 0$  for which we have from (3.56) that  $\mathbf{L} = \pi/2$  and then decreases for increasing  $\tilde{\sigma}_0$ . We can also expand (3.52) to higher orders and obtain

$$\begin{aligned} \frac{\mathcal{F}_{\text{loop}}}{\sqrt{\lambda} k \mathbf{T}} = & -\frac{1}{2} \tilde{\sigma}_0^4 \sqrt{\kappa} + \frac{1}{384} (-128 + 33\pi) \tilde{\sigma}_0^{10} \kappa - \frac{1}{8064} (4112 - 1155\pi) \tilde{\sigma}_0^{16} \kappa^{3/2} \\ & + \frac{(483328 - 99315\pi) \tilde{\sigma}_0^{22}}{5308416} \kappa^2 + \mathcal{O}(\kappa^{5/2}) \quad , \end{aligned} \quad (3.57)$$

which we can use to study phase transitions when comparing it to the Polyakov loop with vanishing free energy. This is depicted in the figures below.

<sup>16</sup>The form of the free energy (3.54) is of a rather different nature when compared with, e.g., (3.36). This suggest that perhaps the correct interpretation of  $\mathbf{L}$  in (3.54) is not as length but instead as a conserved particle number (see Ref. [7] for a discussion of this matter).

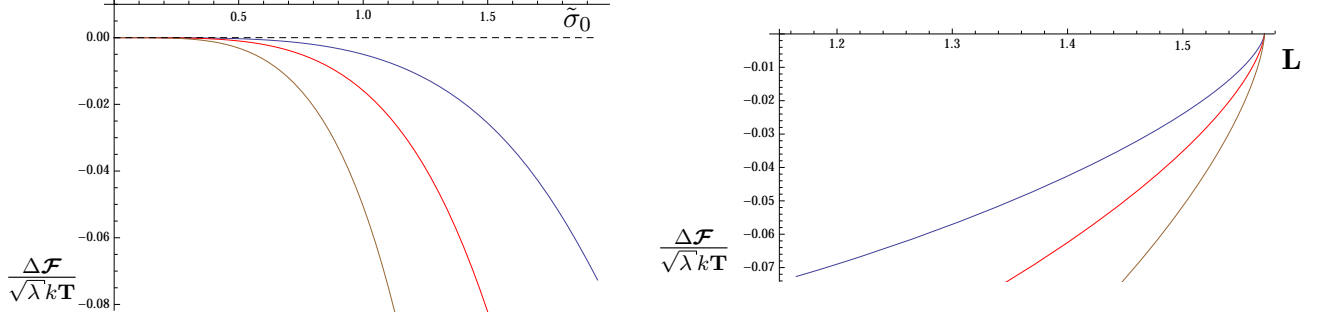


Figure 10: On the left is the plot of  $\frac{\Delta\mathcal{F}}{\sqrt{\lambda}k\mathbf{T}}$  as function of  $\tilde{\sigma}_0$  and as a function of  $\mathbf{L}$  on the right obtained using (3.57) and (3.56). The colour coding is the same as in previous figure. The dashed line is the result for  $\kappa = 0$ .

We can see from the figures below that  $\Delta\mathcal{F} \leq 0$ . Therefore there is no phase transition. As a function of  $\mathbf{L}$ , the free energy decreases for decreasing  $\mathbf{L}$  which is in stark contrast with the previous cases. Since we now have three competing configurations given by (3.41), (3.32) and (3.57) regularised in the same way, one may wonder which is the preferred configuration for a fixed temperature and  $\tilde{\sigma}_0$ . Indeed, as seen in the figures below, the preferred configuration is always the black string with boundary endpoints spacelike separated. On the l.h.s. we have plotted the free energies (3.41) and (3.57) as a function of the bulk depth  $\tilde{\sigma}_0$  while on the r.h.s. we have plotted the difference between (3.41) and (3.57) which we have denoted by  $\Delta\mathcal{F}_{(12)}$ .

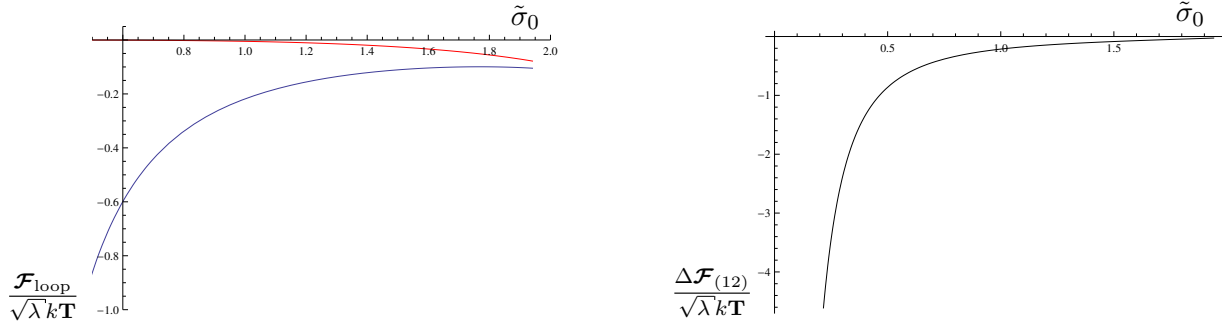


Figure 11: On the left is the plot of  $\mathcal{F}_{\text{loop}}/(\sqrt{\lambda}k\mathbf{T})$  as a function of  $\tilde{\sigma}_0$  for  $\kappa = 0.0001$  for black strings with free energies (3.41) (blue line) and with free energies (3.57) (red line). The qualitative features of this plot are the same for any small value of  $\kappa$ . On the right is the plot of the difference between both, denoted by  $\Delta\mathcal{F}_{12}$ .

From the graph on the l.h.s. we see that the two configurations occupy two different regions of solution space without meeting each other and, in particular, the black strings with spacelike separated boundary endpoints always have lower free energies. From the plot on the r.h.s. we see that the difference  $\Delta\mathcal{F}_{12}$  is always negative, meaning that black strings with free energies (3.41) are always the preferred configurations among the two.

## 4 Black membranes in M-theory

In this section we consider black membranes stretched into the bulk of the M5-brane background ending on one-dimensional strings. At zero temperature, these configurations, known as Wilson surfaces, have been constructed in [14] and their dual interpretation is that of the phase factor associated to the propagation of a very heavy string on the brane. At finite temperature, these solutions have been constructed in [32] using the extremal probe method. Here we will begin by constructing such Wilson surfaces in  $\text{AdS}_7 \times S^4$  using the black probe method, both in the presence of a black hole and without it, and then proceed to construct analogous configurations in thermal  $\text{Sch}_7 \times S^4$ .

### 4.1 Black membranes in Anti-de Sitter

We consider the M-theory background of  $N$  coincident black M5-branes with metric

$$ds^2 = \frac{R^2}{z^2} (-f(z)dt^2 + f(z)^{-1}dz^2 + dx_i^2) + \frac{R^2}{4}d\Omega_{(4)}^2, \quad f(z) = 1 - \gamma \frac{z^6}{z_0^6}, \quad (4.1)$$

where  $i = 1, \dots, 5$  and the horizon radius  $z_0$  is related to the temperature  $T$  via  $z_0 = 3/(2\pi T)$ . The boundary of AdS is located at  $z = 0$ . The parameter  $\gamma$  was introduced as in Sec. 2.2 to distinguish between thermal AdS ( $\gamma = 0$ ) and the AdS black hole ( $\gamma = 1$ ). We now wish to construct a configuration of a stretched black membrane along the  $z$ -direction and whose boundary endpoints are one-dimensional strings placed along the  $x_2$  direction with length  $\mathbb{L}$ . In terms of the world volume coordinates  $(\tau, \sigma, \sigma_2)$ , the embedding map is given by

$$t = \tau, \quad z = \sigma, \quad x_1 = x(\sigma), \quad x_2 = \sigma_2, \quad x_i = 0 \quad \forall i = 3, 4, 5, \quad d\Omega_{(4)} = 0. \quad (4.2)$$

The membrane is stretched from the boundary at  $(x_1, z) = (0, 0)$  to the bulk until  $(x_1, z) = (L/2, \sigma_0)$  and back again to the boundary at  $(x_1, z) = (L, 0)$ . The problem is symmetric around  $x_1 = L/2$  and also translational invariant along  $x_2$  so this direction can be integrated out and plays no role in the problem. The induced metric becomes

$$\gamma_{ab}d\sigma^a d\sigma^b = \frac{R^2}{\sigma^2} (-f(\sigma)d\tau^2 + (f(\sigma)^{-1} + x'(\sigma)^2)d\sigma^2 + d\sigma_2^2), \quad (4.3)$$

while the norm of the timelike Killing vector field on the world volume is  $\mathbf{k} = R\sqrt{f}/\sigma$ . The probe black membrane being stretched carries a  $Q_2$  charge, which in  $D = 11$  requires  $n = 6$  in Eqs.(2.5)-(2.7). The free energy (2.12) can then be written as

$$\mathcal{F}[x(\sigma)] = -2AL \left( \frac{3}{2\pi T} \right)^6 \int_0^{\sigma_0} d\sigma \sqrt{1 + x'(\sigma)^2} F(\sigma), \quad F(\sigma) = \frac{R^9}{\sigma^9} f(\sigma)^3 \frac{1 + 6 \sinh^2 \alpha}{\cosh^6 \alpha}, \quad (4.4)$$

where we have integrated along the direction  $\sigma_2$  yielding the overall factor of  $\mathbb{L}$ . The solution to the equations of motion that follow from (4.15) is identical to (3.21) but with the function  $G(\sigma)$  replaced by  $F(\sigma)$ , i.e.,

$$x'(\sigma) = \left( \frac{f(\sigma)^2 F(\sigma)^2}{f(\sigma_0) F(\sigma_0)^2} - f(\sigma) \right)^{-\frac{1}{2}}, \quad (4.5)$$

where we have imposed  $x'(\sigma) \rightarrow \infty$  as  $\sigma \rightarrow \sigma_0$ . The boundary conditions (2.2) are also satisfied. We also note that the equations of motion admit a solution of the type  $x'(\sigma) = 0$  corresponding to a straight membrane stretched from the boundary to the bulk, analogous to the Polyakov loop in Sec. 2.2.

## Critical distance and regime of validity

Introducing the dimensionless coordinate  $\hat{\sigma} = 2\pi T\sigma/3$ , one can define a parameter  $\bar{\kappa}$  as in (2.23) via

$$\bar{\kappa} \equiv \frac{Q_2}{6AR^6} = \frac{f(\hat{\sigma})^3}{\hat{\sigma}^6} \frac{\sinh \alpha(\hat{\sigma})}{\cosh^5 \alpha(\hat{\sigma})} , \quad (4.6)$$

where  $f(\hat{\sigma}) = 1 - \gamma\hat{\sigma}^6$ . The fact that the function  $f(\hat{\sigma})$  depends on a different power of  $\hat{\sigma}$  implies that the critical distance will change. In particular, we have that

$$\hat{\sigma}_c^2|_{\gamma=1} = \frac{\left(18 + \sqrt{324 + 75\sqrt{5}\bar{\kappa}}\right)^{2/3} - 3^{1/3}5^{5/6}\bar{\kappa}^{1/3}}{6^{2/3}\left(18 + \sqrt{324 + 75\sqrt{5}\bar{\kappa}}\right)^{1/3}} , \quad (4.7)$$

while for  $\gamma = 0$  we obtain again (2.25). We see that the membranes in the case  $\gamma = 1$  cannot stretch all the way to the horizon for non-zero  $\bar{\kappa}$ , i.e.,  $\hat{\sigma}_c \leq 1$ . These solutions, however, are not valid all the way to  $\hat{\sigma}_c$  as we will see below.

The validity of the approximation requires that one can ignore quantum corrections. The largest scale characterising the membrane is  $r_c = R\bar{\kappa}^{1/6}$ . We therefore need to require  $r_c \gg l_p$  which implies that  $N_{(2)} \gg 1$  where  $N_{(2)}$  stands for the number of M2 black probe membranes and where we have used the fact that the brane tension can be written as  $T_{(p)} = ((2\pi)^p l_p^{p+1})^{-1}$ . Requiring the radius  $r_c$  to be much smaller than the curvature of the background  $r_c \ll R$  yields the condition  $N_{(2)} \ll N$  where  $N$  should now be seen as the number of M5-branes that compose the background geometry.<sup>17</sup> Therefore we must have that

$$1 \ll N_{(2)} \ll N . \quad (4.8)$$

Note that in terms of these quantities, we can write  $\bar{\kappa}$ , defined in (4.6), as  $\bar{\kappa} = N_{(2)}/(2N^2)$ . The requirement  $r_c \ll R$  and  $r_c \ll 1/T$  lead again to (2.29). Furthermore, we should have small variations of the local temperature compared to the background temperature,

$$r_c \frac{\mathcal{T}'}{T} = \frac{r_c}{R\sqrt{f}} + 3\gamma r_c \frac{(2\pi T\sigma/3)^6}{Rf^{3/2}} \ll 1 . \quad (4.9)$$

Near the boundary this is satisfied as well as in general for the case  $\gamma = 0$  given that  $\bar{\kappa} \ll 1$ . In the case  $\gamma = 1$  and near the critical distance we find that the l.h.s. above is proportional to  $\bar{\kappa}^{-1/3}$  and hence the approximation breaks down near  $\sigma_c$ . For distances close to  $z_0$  we need to require  $z_0 - \sigma \gg \bar{\kappa}^{1/9}z_0$ . The extrinsic curvature length takes the same form as in (2.31) and the requirement  $r_c \ll L_{\text{ext}}(\sigma)$  is satisfied provided  $z_0 - \sigma \gg \bar{\kappa}^{1/9}z_0$ .

## Regularized free energy and large $N$ expansion

The free energy (4.15) can be rewritten as

$$\mathcal{F}[x(\hat{\sigma})] = \frac{4^2}{3^2} \pi \mathbb{L} N_{(2)}^{3/2} \sqrt{\lambda_M} T^2 \int_0^{\hat{\sigma}_0} \frac{d\hat{\sigma}}{\hat{\sigma}^3} (1 - X) \sqrt{1 + f(\hat{\sigma})x'(\hat{\sigma})^2} , \quad (4.10)$$

where we have introduced the 't Hooft-like coupling  $\lambda_M = N^2/N_{(2)}$  found in [30] in the context of the M2-M5 intersection. The characteristic  $N_{(2)}^{3/2}$  dependence of M2-branes is also found in the context of

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<sup>17</sup>Here we have used the AdS/CFT dictionary  $R^3 = 2N/(\pi T_{(2)})$ .

giant gravitons in this background [31]. The free energy diverges at  $\hat{\sigma} = 0$  so, as in the previous cases, we must subtract a piece of the form

$$\mathcal{F}_{\text{sub}} = \frac{4^2}{3^2} \pi \mathbb{L} N_{(2)}^{3/2} \sqrt{\lambda_M} T^2 \int_0^{\hat{\sigma}_{\text{cut}}} \frac{d\hat{\sigma}}{\hat{\sigma}^3} (1 - X) \ , \quad (4.11)$$

such that  $\Delta \mathcal{F} = \mathcal{F} - \mathcal{F}_{\text{sub}} = \mathcal{F}_{\text{surf}} - 2\mathcal{F}_{\text{P}_s}$ , where

$$\mathcal{F}_{\text{surf}} = \frac{4^2}{3^2} \pi \mathbb{L} N_{(2)}^{3/2} \sqrt{\lambda_M} T^2 \left( -\frac{1}{2\hat{\sigma}_0^2} + \int_0^{\hat{\sigma}_0} \frac{d\hat{\sigma}}{\hat{\sigma}^3} \left( (1 - X) \sqrt{1 + f x'(\sigma)^2} - 1 \right) \right) \ , \quad (4.12)$$

$$\mathcal{F}_{\text{P}_s} = -\frac{1}{2} \frac{4^2}{3^2} \pi \mathbb{L} N_{(2)}^{3/2} \sqrt{\lambda_M} T^2 \left( \frac{1}{2\hat{\sigma}_{\text{cut}}^2} + \int_0^{\hat{\sigma}_{\text{cut}}} \frac{d\hat{\sigma}}{\hat{\sigma}^3} X \right) \ . \quad (4.13)$$

Here  $\mathcal{F}_{\text{surf}}$  is the regularised free energy of the stretched black membrane while  $\mathcal{F}_{\text{P}_s}$  is the regularised free energy of the straight black membrane. By similar arguments as those used for the Polyakov loop, but now requiring configurations at  $T = 0$  for any  $\bar{\kappa}$  to have the same value of  $\partial S / \partial T$ , one finds that  $\mathcal{F}_{\text{P}_s} = -(4^2/3^2) \gamma \pi \mathbb{L} N_{(2)}^{3/2} \sqrt{\lambda_M} T^2 / 4$ . We can now proceed and obtain analytic results in a small  $\bar{\kappa}$  expansion. We find that the length along the direction  $x_1$  for small  $\bar{\kappa}$  and small  $\hat{\sigma}_0$  is given by

$$LT = \frac{3\Gamma(\frac{2}{3})}{\sqrt{\pi}\Gamma(\frac{1}{6})} \hat{\sigma}_0 + \frac{1}{3\sqrt{\pi}} \left( \frac{\Gamma(\frac{2}{3})}{\Gamma(\frac{1}{6})} + \frac{\Gamma(\frac{7}{6})}{\Gamma(-\frac{1}{3})} \right) \hat{\sigma}_0^4 \sqrt{\bar{\kappa}} - \gamma \frac{27\Gamma(\frac{5}{3})}{28\sqrt{\pi}\Gamma(\frac{1}{6})} \hat{\sigma}_0^7 + \mathcal{O}(\bar{\kappa}) + \mathcal{O}(\hat{\sigma}_0^8) \ . \quad (4.14)$$

Again, we see that the leading order correction only depends on the properties of the probe and not of the background. The background contribution when  $\gamma = 1$  had already been obtained in [32]. Expanding the free energy (4.12) for small  $\kappa$  and  $L$  and inverting (4.14) we find

$$\begin{aligned} \mathcal{F}_{\text{surf}} = & -\frac{4^2 \pi}{3^2} \frac{\mathbb{L}}{L^2} N_{(2)}^{3/2} \sqrt{\lambda_M} \left( \frac{3^2}{2\sqrt{\pi}} \frac{\Gamma(\frac{2}{3})^3}{\Gamma(\frac{1}{6})^3} - \frac{2\Gamma(-\frac{2}{3})\Gamma(\frac{7}{6})^2}{3\sqrt{3}\Gamma(\frac{2}{3})} \sqrt{\bar{\kappa}} (TL)^3 \right. \\ & \left. + \gamma \frac{2\pi^{5/2}}{567} \frac{\Gamma(\frac{1}{6})^3}{\Gamma(\frac{2}{3})^3} (TL)^6 + \mathcal{O}(\bar{\kappa}) + \mathcal{O}(L^8) \right) \ . \end{aligned} \quad (4.15)$$

The leading order result agrees with the result presented in [14] and the coefficient proportional to  $\gamma$  agrees with the one found in [32] when  $\gamma = 1$ . Again, the corrections due to the probe are the leading order corrections and they enter at the same order as in the case of the Wilson loops of Sec. 2.2. On the other hand the contribution due to the background enters at higher order when compared to the Wilson loops of Sec. 2.2. For the black probe corrections to be leading it is sufficient to require

$$(TL)^3 \ll \frac{1}{\lambda_M} \ , \quad (4.16)$$

where we have used the fact that, as noted above,  $\bar{\kappa} = N_{(2)}/(2N^2)$ . This is always satisfied for small enough temperatures, and more easily satisfied for small temperatures than in the Wilson loop case due to the cubic factor in  $TL$ . The combination  $\mathcal{F}_{\text{surf}} L^2 / \mathbb{L}$  is scale invariant as well as the combination  $LT$ , therefore we see the underlying conformal symmetry of the dual gauge theory.

## Numerical analysis

We now analyse the solution space in detail by solving numerically for the various quantities. We begin by solving for the distance between the strings on the boundary for both values of  $\gamma$ . This is depicted in the figures below.

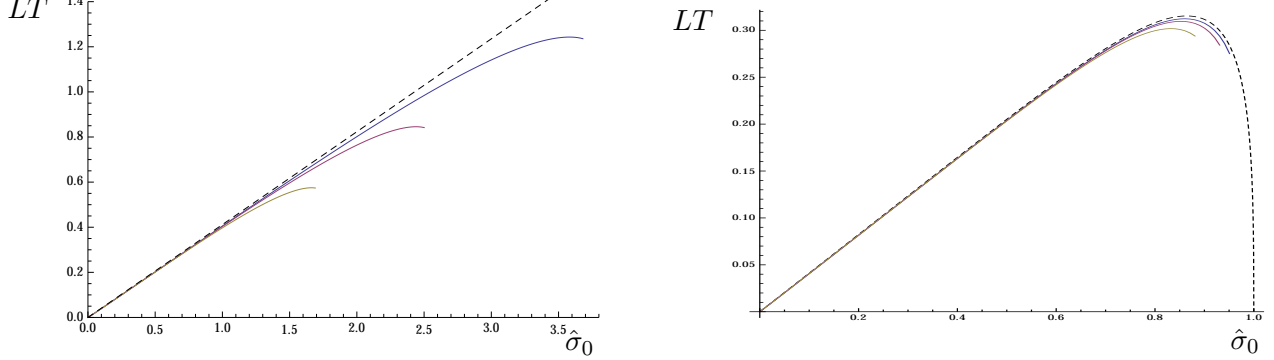


Figure 12: On the left we have the numerical results for the dimensionless length  $LT$  as a function of  $\hat{\sigma}_0$  in thermal AdS for  $\bar{\kappa} = 0.01$  (yellow curve),  $\bar{\kappa} = 0.001$  (red curve) and  $\bar{\kappa} = 0.0001$  (blue curve). The dashed line is the result for  $\bar{\kappa} = 0$ . On the right we have the numerical results in the presence of a black hole.

The qualitative features are the same as for the Wilson loop of Sec. 2.2. In particular we observe the existence of the critical distance  $\hat{\sigma}_c$  beyond which solutions cease to exist. In the presence of the black hole, this means that the black membrane cannot stretch all the way to the horizon. The dashed curve representing the case  $\kappa = 0$  can be obtained exactly,

$$LT|_{\kappa=0} = \frac{3\Gamma(\frac{2}{3})}{\sqrt{\pi}\Gamma(\frac{1}{6})} \hat{\sigma}_0 \sqrt{1 - \gamma \hat{\sigma}_0^6} {}_2F_1\left(\frac{1}{2}, \frac{2}{3}, \frac{7}{6}; \gamma \hat{\sigma}_0^6\right), \quad (4.17)$$

and agrees with the result of [32] when  $\gamma = 1$ . We now present the free energy difference between the two configurations. On the l.h.s. of the plot below we have the dependence of the free energy difference as a function of  $\hat{\sigma}_0$  and on the r.h.s. as a function of  $LT$  in the case  $\gamma = 1$ . We observe that  $\Delta\mathcal{F}$  becomes positive for large values of  $LT$ , which implies a similar effect as the Debye screening effect for the quark-antiquark pair. The exact onset can be obtained numerically and reads

$$LT|_{\Delta\mathcal{F}=0} \sim 0.2779 + 0.1131\sqrt{\bar{\kappa}}. \quad (4.18)$$

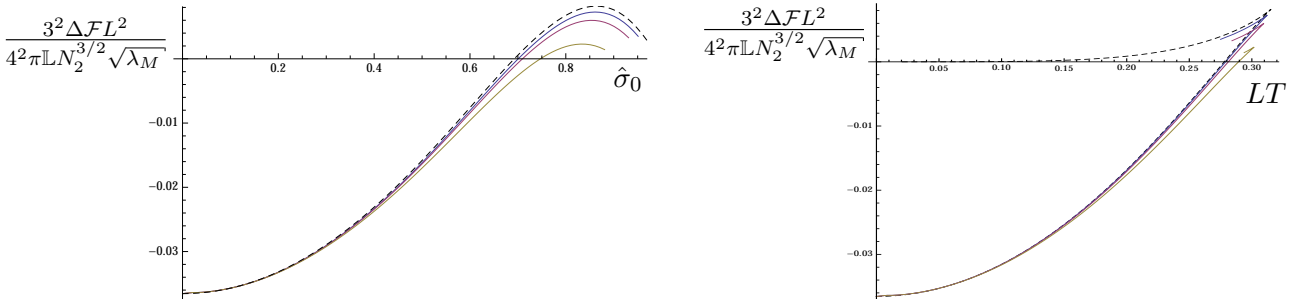


Figure 13: On the left we have the free energy difference as a function of  $\hat{\sigma}_0$  and on the right as a function of  $LT$  in the case  $\gamma = 1$ . The dashed line represents  $\bar{\kappa} = 0$  and has been obtained in [32].



In the case  $\gamma = 0$  the results are also similar to the Wilson loop as one can see in the figures below, namely, there is no phase transition.

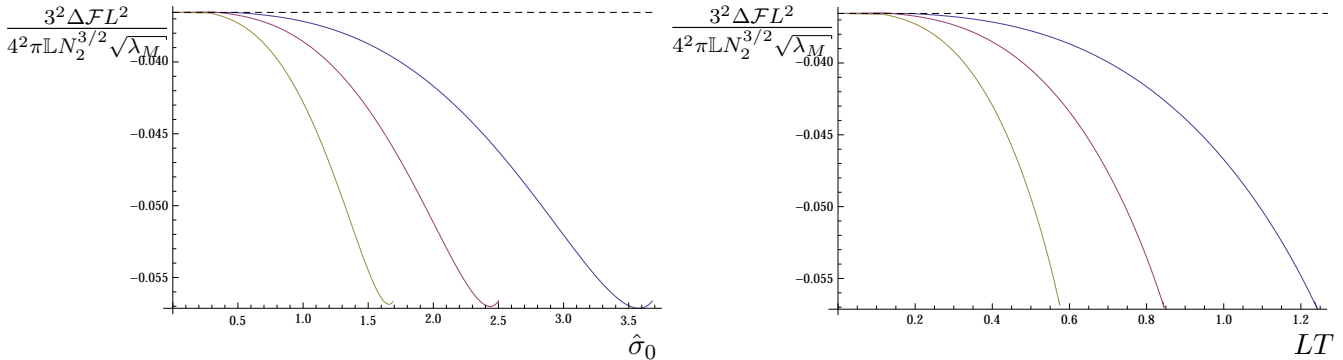


Figure 14: On the left we have the free energy difference as a function of  $\hat{\sigma}_0$  and on the right as a function of  $LT$  for  $\gamma = 0$ . The colour coding is the same as in the previous figure.

## 4.2 Black membranes in Schrödinger

In this section we construct several membrane geometries in  $\text{Sch}_7 \times S^4$ , obtained via the null Melvin twist of the extremal M5-brane. This spacetime geometry can be written as [16]

$$ds^2 = \frac{R^2}{z^2} \left( -\frac{\ell^4}{z^4} dt^2 + 2dt d\xi + dz^2 + dx_i^2 \right) + \frac{R^2}{4} d\Omega_{(4)}^2, \quad (4.19)$$

where  $i = 1, \dots, 4$  and  $\ell^2 = \beta R^3$  measures the deformation of null-AdS into Schrödinger.<sup>18</sup> As in the case of the  $\text{Sch}_5 \times S^5$  introduced in (3.1), boosts in the  $(t, \xi)$  plane or rescalings of the  $\text{Sch}_7$  coordinates imply that the physical quantities describing configurations at finite temperature and with characteristic lengths  $L$  and  $\mathbb{L}$  are given in terms of the dimensionless quantities

$$\mathbf{T} = T\ell, \quad \mathbf{L} \equiv \frac{L}{\ell}, \quad \tilde{\mathbb{L}} \equiv \frac{\mathbb{L}}{\ell}. \quad (4.20)$$

In this spacetime we will construct four different types of stretched black membranes. The first configuration we construct is the analogue of the one presented in Sec. 3.1 and consists of an interpolating membrane between the bulk dual to the AdS Wilson surface presented in the previous section and a new Schrödinger configuration. The boundary endpoints of these configurations are one-dimensional strings with length  $\mathbb{L}$  and the remaining three configurations are differentiated depending on how these strings are placed: (i) spatially extended and spacelike separated, (ii) null extended and spacelike separated, and (iii) spatially extended and lightlike separated.

Before analysing each configuration individually, we note that all these last three configurations share the same form of  $\mathbf{k} = R\ell^2/z^3$ , therefore allowing us to make general statements about their validity. In particular, we can define the parameter  $\bar{\kappa}$  as in (4.6) via

$$\bar{\kappa} = \frac{1}{\tilde{\sigma}^{18}} \frac{\sinh \alpha(\tilde{\sigma})}{\cosh^5 \alpha(\tilde{\sigma})}, \quad \tilde{\sigma} = \left( \frac{2\pi T}{3 \ell^2} \right)^{\frac{1}{3}} \sigma, \quad (4.21)$$

<sup>18</sup>Note that our parameter  $\beta$  has been rescaled when compared with the parameter  $\beta$  used in [16], namely, we have taken the metric in [16] and performed the rescaling  $\beta \rightarrow \beta\sqrt{2}$ .

where the rescaling of the coordinate  $\tilde{\sigma}$  ensures that  $\bar{\kappa}$  is the same both in AdS and in Schrödinger. With this we can obtain the critical distance  $\tilde{\sigma}_c$  beyond which these three black membranes in this spacetime cannot extend

$$\tilde{\sigma}_c^2 = \frac{2^{\frac{4}{9}}}{5^{\frac{5}{18}}} \bar{\kappa}^{-\frac{1}{9}} . \quad (4.22)$$

The remaining validity analysis of these configurations is similar to the case  $\gamma = 0$  studied in the previous section. One should, however, pay attention to the variation of the local temperature  $r_c \mathcal{T}'/T = 3r_c \sigma^2/(R\ell^2) \ll 1$  which implies near the critical distance that

$$\kappa \ll \left(\frac{\ell}{T}\right)^{12} . \quad (4.23)$$

Again, this requirement is generic for all the three types of black membranes mentioned above and imposes a mild bound on the ratio  $\ell/T$ . Furthermore, a careful analysis of the extrinsic curvature length  $L_{\text{ext}}(\sigma)$ , which can be evaluated explicitly using the information provided below for each black membrane, tell us that no new requirements appear. This indicates that these configurations are valid solutions all the way up to  $\sigma_c$ . We will now analyse each of these three cases and also the case of the interpolating membrane one-by-one but we will be succinct in the details since the method we follow has been repeatedly used in previous sections.

### Interpolating membranes with spatially extended and spacelike separated boundary strings

In this section we construct a membrane configuration which interpolates between the Wilson surface analysed in the previous section, with  $\gamma = 0$ , and a new Schrödinger configuration when  $\ell$  is large, which we will analyse separately in the next section. These are the analogue of the interpolating black strings studied in Sec. 3.1. We begin by introducing the coordinates  $(t', \xi')$  as in (3.4) which brings the metric (4.19) to the form.

$$ds^2 = \frac{R^2}{z^2} \left[ -y(z) dt'^2 - \frac{\ell^4}{z^4} dt' d\xi' + \left(1 - \frac{\ell^4}{2z^4}\right) d\xi'^2 + dz^2 + dx_i^2 \right] + R^2 d\Omega_{(4)}^2 , \quad g(z) = 1 + \frac{\ell^4}{2z^4} . \quad (4.24)$$

By taking  $\ell \rightarrow 0$  we obtain the metric (4.1) with  $\gamma = 0$  where  $t'$  is identified with  $t$  and  $\xi'$  with  $x_3$ . We now choose an embedding map such that

$$t' = \tau , \quad z = \sigma , \quad x_1 = x(\sigma) , \quad x_2 = \sigma_2 , \quad \xi' = 0 , \quad x_i = 0 , \forall i = 3, 4 , \quad d\Omega_{(4)} = 0 , \quad (4.25)$$

which results in an induced metric of the form

$$\gamma_{ab} d\sigma^a d\sigma^b = \frac{R^2}{\sigma^2} \left( -y(\sigma) d\tau^2 + (1 + x'(\sigma)^2) d\sigma^2 + d\sigma_2^2 \right) . \quad (4.26)$$

This metric reduces to (4.3) when  $\ell = 0$ . The pullback of the timelike Killing vector field  $\partial_{t'}$  onto the world volume  $\mathcal{W}_3$  gives rise to  $\mathbf{k} = (R/\sigma)\sqrt{y(\sigma)}$ . Hence, if  $\ell = 0$  we obtain  $\mathbf{k} = R/\sigma$  which is the result obtained in AdS when  $\gamma = 0$ , while if we take  $\ell \rightarrow \infty$  we get  $\mathbf{k} \sim R\ell^2/\sigma^3$ .

Solving now the equation of motion that arises from varying (2.12) leads to the familiar solution

$$x'(\sigma) = \left( \frac{U(\sigma)^2}{U(\sigma_0)^2} - 1 \right)^{-\frac{1}{2}} , \quad U(\sigma) = \frac{R^9}{\sigma^9} y(\sigma)^{\frac{7}{2}} \frac{1 + 6 \sinh^2 \alpha}{\cosh^6 \alpha} . \quad (4.27)$$

We now focus on the zero-temperature case with  $\kappa = 0$ . In order to do so we consider the regularised free energy which takes the form

$$\mathcal{F}_{\text{surf}} = \frac{4\mathbb{L}N_{(2)}^{3/2}}{\pi}\sqrt{\lambda_M}\left(u(\sigma_0) + \int_0^{\sigma_0} \frac{d\sigma}{\sigma^3}\sqrt{y(\sigma)}\left((1-X)\sqrt{1-x'(\sigma)^2}-1\right)\right), \quad (4.28)$$

where we have defined the function  $u(\sigma_0)$  via

$$u(\sigma_0) = -\frac{1}{4}\left(\frac{\sqrt{\ell^4 + 2\sigma_0^4}}{\sqrt{2}\sigma_0^4} + \sqrt{2}\frac{\text{arcsinh}(\frac{\ell^2}{\sqrt{2}\sigma_0^2})}{\ell^2}\right). \quad (4.29)$$

The Polyakov surface can be argued to have zero free energy for any value of  $\ell, \kappa, T$  with similar arguments as in the previous cases. We now obtain expressions for the length  $L$  and  $\mathcal{F}_{\text{surf}}$  as a function of  $\sigma_0$  at  $T = 0$ . This behaviour is shown in the plots below.

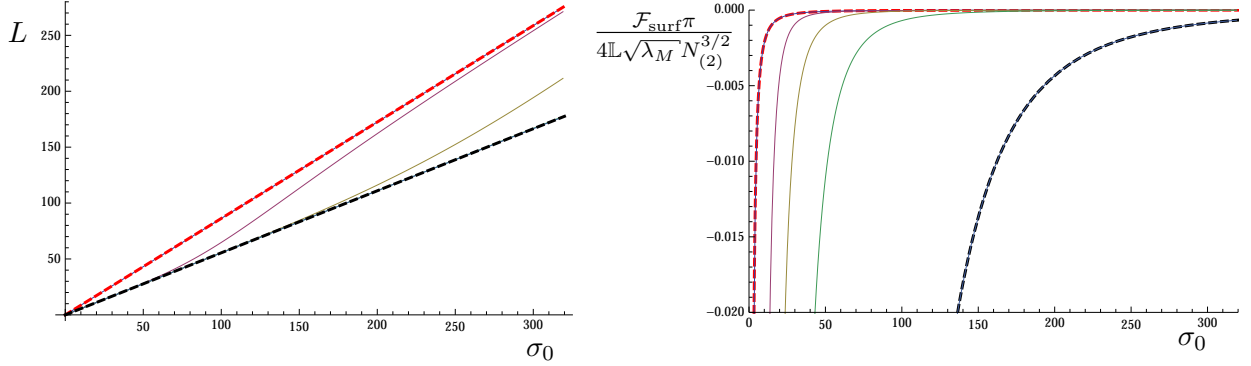


Figure 15: On the left we display the behaviour of the distance  $L$  as function of  $\sigma_0$  at  $T = 0$ . The dashed solid red line corresponds to the AdS result given by the first term in (4.14) while the dashed solid black line corresponds to the result (4.30). The blue line represents  $(\ell/2^{1/4}) = 1$ , the magenta line  $(\ell/2^{1/4}) = 100$ , the yellow line  $(\ell/2^{1/4}) = 300$ , the green line  $(\ell/2^{1/4}) = 1000$  and the cyan line  $(\ell/2^{1/4}) = 10000$ . On the right is the behaviour of the free energy as a function of  $\sigma_0$  at  $T = 0$ . The dashed solid red line is the first term obtained in (4.12) while the dash solid black line is the result obtained in (4.30).

As one can see, the behaviour exhibited in the above plots resembles the case of black strings in Sec. 3.1. The red dashed lines correspond to the leading order results for the length and free energy obtained in (4.14) and (4.12) respectively. The black dashed lines correspond to the results which will be obtained in the next section and to leading order read

$$L|_{\ell \rightarrow \infty} = 2\sqrt{\pi}\sigma_0 \frac{\Gamma(\frac{3}{5})}{\Gamma(\frac{1}{10})}, \quad \frac{\mathcal{F}_{\text{surf}}\pi}{4\mathbb{L}\sqrt{\lambda_M}N_{(2)}^{3/2}}|_{\ell \rightarrow \infty} = \frac{\sqrt{\pi}\ell^2}{10\sigma_0^4} \frac{\Gamma(-\frac{2}{5})}{\Gamma(\frac{1}{10})}. \quad (4.30)$$

In particular, values of the length  $L$  and the free energy are bound between the AdS values and a purely Schrödinger configuration which will be analysed in the next section. Furthermore, we observe that  $\mathcal{F}_{\text{surf}}$  is proportional to  $L^{-2}$  for small  $\ell$  and to  $L^{-4}$  for large  $\ell$ , since the relation between  $L$  and  $\sigma_0$  is roughly linear.

We can now consider the finite temperature versions of these geometries. However, all the qualitative features are the same as those presented in Sec. 3.1. We will therefore only consider in detail the heated

up version of the configuration in the limit  $\ell \rightarrow \infty$  in the next section. Nevertheless, we note that at finite temperature we can define a parameter  $\hat{\ell}$  such that

$$y(\hat{\sigma}) = 1 + \frac{\hat{\ell}^4}{\hat{\sigma}^4} \quad , \quad \hat{\ell}^4 = \frac{2^3 \pi^4 T^4 \ell^4}{3^4} \quad . \quad (4.31)$$

In terms of this parameter and of the parameter  $\bar{\kappa}$  defined (4.6) we can find the critical distance for these configurations, however, its form is too cumbersome to be presented here. The resulting behaviour of the free energy is similar to what is presented in Fig. 6, namely, the solution terminates for a certain value of  $\hat{\sigma}_c$ .

### Membranes with spatially extended and spacelike separated boundary strings

In this section we construct the membrane geometry which in the limit  $\ell \rightarrow \infty$  connects with the interpolating membrane solution found in the previous section. In order to do so, we choose the embedding map

$$t = \tau \quad , \quad z = \sigma \quad , \quad x_1 = x(\sigma) \quad , \quad x_2 = \sigma_2 \quad , \quad \xi = x_i = 0 \quad , \quad \forall i = 3, 4 \quad , \quad d\Omega_{(4)} = 0 \quad , \quad (4.32)$$

which is the analogous configuration in  $\text{AdS}_7 \times S^4$  to the one constructed in Sec. 3.2. The world volume inherits the metric

$$\gamma_{ab} d\sigma^a d\sigma^b = \frac{R^2}{\sigma^2} \left( -\frac{\ell^4}{\sigma^4} d\tau^2 + (1 + x'(\sigma)^2) d\sigma^2 + d\sigma_2^2 \right) \quad , \quad (4.33)$$

which when introduced in (2.12) and after careful variation leads to the solution

$$x'(\sigma) = \left( \frac{K(\sigma)^2}{K(\sigma_0)^2} - 1 \right)^{\frac{1}{2}} \quad , \quad K(\sigma) = \frac{R^9 \ell^{14}}{\sigma^{23}} \frac{1 + 6 \sinh^2 \alpha}{\cosh^6 \alpha} \quad , \quad (4.34)$$

where we have imposed the same boundary conditions as in the previous section. The equations of motion, as usual, also allow for Polyakov surfaces with  $x'(\sigma) = 0$ .

Since these configurations are valid only for small  $\bar{\kappa}$  we proceed and analyse their properties in a small  $\bar{\kappa}$  expansion. We consider first the distance between the two strings on the boundary. With the help of the expansion

$$\cosh^2 \alpha = \frac{1}{\tilde{\sigma}^9 \sqrt{\bar{\kappa}}} - \frac{1}{4} - \frac{5}{32} \tilde{\sigma}^9 \sqrt{\bar{\kappa}} + \mathcal{O}(\bar{\kappa}) \quad , \quad (4.35)$$

we find, to leading order in the bulk depth  $\tilde{\sigma}_0$ , the dimensionless distance

$$\mathbf{L} \mathbf{T}^{\frac{1}{3}} \left( \frac{\pi}{12} \right)^{\frac{1}{3}} = \frac{\sqrt{\pi} \Gamma(\frac{3}{5})}{\Gamma(\frac{1}{10})} \tilde{\sigma}_0 + \frac{\pi}{30} \left( \frac{2\Gamma(\frac{3}{5})}{\sqrt{\pi} \Gamma(\frac{1}{10})} - 1 \right) \tilde{\sigma}_0^{10} \sqrt{\bar{\kappa}} + \mathcal{O}(\bar{\kappa}) \quad . \quad (4.36)$$

The leading order term, when written in terms of  $\sigma_0$  leads to the result presented in (4.30). We now take a look at the free energy. After implementing a similar regularisation scheme as in the previous sections we arrive at the regularised free energy

$$\mathcal{F}_{(1)}[x(\tilde{\sigma})] = c_n \tilde{\mathbf{L}} N_{(2)}^{\frac{3}{2}} \sqrt{\lambda_M} \mathbf{T}^{\frac{4}{3}} \left( -\frac{1}{4\tilde{\sigma}_0^4} + \int_0^{\tilde{\sigma}_0} \frac{d\tilde{\sigma}}{\tilde{\sigma}^5} \left( (1 - X) \sqrt{1 + x'(\tilde{\sigma})^2} - 1 \right) \right) \quad , \quad (4.37)$$

where we have defined the constant  $c_n = (2^{10} \pi / 3^4)^{1/3}$  and introduced the dimensionless quantities defined in (4.20). The Polyakov surface, corresponding to the solution  $x'(\sigma) = 0$ , can be argued, following the

same reasoning as the previous section, to have zero free energy. Expanding now in small  $\bar{\kappa}$  and small distance  $L$ , and inverting (4.36), we find

$$\mathcal{F}_{(1)} = c_n \frac{\tilde{\mathbb{L}}}{\mathbf{L}^4} N_{(2)}^{\frac{3}{2}} \sqrt{\lambda_M} \left( \frac{3^{4/3} \pi^{\frac{7}{6}} \Gamma(-\frac{2}{5}) \Gamma(\frac{3}{5})^4}{2^{\frac{1}{3}} 12500 \Gamma(\frac{1}{10}) \Gamma(\frac{11}{10})^4} + C_1 \sqrt{\bar{\kappa}} \mathbf{L}^9 \mathbf{T}^3 \right) + \mathcal{O}(\bar{\kappa}) + \mathcal{O}(\mathbf{L}^{10}) , \quad (4.38)$$

where  $C_1$  is a real numerical constant which we give in App. A. The leading order term in this equation, when written in terms of  $\sigma_0$  leads to the result presented in (4.30). From (4.38) we can see that the combination  $\mathcal{F}_{(1)} \mathbf{L}^4 / (\tilde{\mathbb{L}})$  is scale invariant as well as  $\mathbf{L} \mathbf{T}^{1/3}$ , hence exhibiting the scale invariance of the dual gauge theory. We can also depict the entire solution space. It exhibits the same qualitative features as those encountered in Sec. 3.2 for black strings in Schrödinger with spacelike separated boundary endpoints. In particular there is no phase transition and this configuration is always the preferred one with respect to the Polyakov surface. We provide details in App. A.

### Membranes with null extended and spacelike separated boundary strings

In order to construct these configurations we choose the embedding map

$$t = \tau , \quad z = \sigma , \quad x_1 = x(\sigma) , \quad \xi = \sigma_2 , \quad x_i = 0 , \quad \forall i = 2, 3, 4 , \quad d\Omega_{(4)} = 0 , \quad (4.39)$$

leading to the stationary induced metric

$$\gamma_{ab} d\sigma^a d\sigma^b = \frac{R^2}{\sigma^2} \left( -\frac{\ell^4}{\sigma^4} d\tau^2 + 2d\tau d\sigma_2 + (1 + x'(\sigma)^2) d\sigma^2 \right) . \quad (4.40)$$

The boundary strings now lie in the null  $\xi$ -direction. Since these configurations are translational invariant along  $\xi$ , this direction does not play a role and can be integrated out leading to an overall factor of  $\mathbb{L}$  in the free energy. Using the induced metric (4.40) and solving the equation of motion obtained by varying (2.12) we obtain the solution

$$x'(\sigma) = \left( \frac{Y(\sigma)^2}{Y(\sigma_0)^2} - 1 \right)^{\frac{1}{2}} , \quad Y(\sigma) = \frac{R^9 \ell^{12}}{\sigma^{21}} \frac{1 + 6 \sinh^2 \alpha}{\cosh^6 \alpha} . \quad (4.41)$$

We now study this solution in the small  $\bar{\kappa}$  regime using the expansion (4.35). To leading order in  $\bar{\kappa}$  we find the dimensionless length

$$\mathbf{L} \mathbf{T}^{\frac{1}{3}} \left( \frac{\pi}{12} \right)^{\frac{1}{3}} = \frac{\sqrt{\pi} \Gamma(\frac{2}{3})}{\Gamma(\frac{1}{6})} \tilde{\sigma}_0 + \frac{\Gamma(\frac{2}{3}) \Gamma(\frac{5}{6}) + \sqrt{3} \Gamma(-\frac{2}{3}) \Gamma(\frac{13}{6})}{18\sqrt{\pi}} \sqrt{\bar{\kappa}} \tilde{\sigma}_0^{10} + \mathcal{O}(\bar{\kappa}) . \quad (4.42)$$

The regularised free energy in turn reads

$$\mathcal{F}_{(2)}[x(\tilde{\sigma})] = c_s \tilde{\mathbb{L}} N_{(2)}^{\frac{3}{2}} \sqrt{\lambda_M} \mathbf{T}^{\frac{2}{3}} \left( -\frac{1}{2\tilde{\sigma}_0^2} + \int_0^{\tilde{\sigma}_0} \frac{d\tilde{\sigma}}{\tilde{\sigma}^3} \left( (1 - X) \sqrt{1 + x'(\tilde{\sigma})^2} - 1 \right) \right) , \quad (4.43)$$

where we have defined the constant  $c_s = (2^8/(3^2\pi))^{1/3}$ . Note that the background subtraction here is different than the one used to regularise (4.37). Using the expansion (4.35) and inverting (4.42) we obtain

$$\mathcal{F}_{(2)} = c_s \frac{\tilde{\mathbb{L}}}{\mathbf{L}^2} N_{(2)}^{\frac{3}{2}} \sqrt{\lambda_M} \left( \frac{2^{\frac{1}{3}} \Gamma(-\frac{1}{3})^3}{3^{\frac{1}{3}} 9 \Gamma(\frac{1}{6})^3} + C_2 \sqrt{\bar{\kappa}} \mathbf{L}^9 \mathbf{T}^3 \right) + \mathcal{O}(\bar{\kappa}) + \mathcal{O}(L^{10}) , \quad (4.44)$$

where  $C_2$  is a real numerical constant that we give in App. A. The general form of  $\mathcal{F}_{(2)}$  is in accordance with scale invariance of the dual theory. The full solution space of these configurations is also very similar to the black strings constructed in Sec. 3.2 and hence we leave these details to App. A. However, we note that there is no phase transition with respect to the Polyakov surface and that this is always the preferred configuration. Also note that since we have made a different regularisation procedure in (4.37) and (4.43) what is meant here by Polyakov surface is also different for both configurations.

### Membranes with spatially extended and lightlike separated boundary strings

The last configuration we construct in this section is obtained by choosing the embedding map

$$t = \tau \quad , \quad z = \sigma \quad , \quad \xi = \xi(\sigma) \quad , \quad x_1 = \sigma_2 \quad , \quad x_i = 0 \quad , \quad \forall i = 2, 3, 4 \quad , \quad d\Omega_{(4)} = 0 \quad , \quad (4.45)$$

which leads to the induced metric

$$\gamma_{ab} d\sigma^a d\sigma^b = \frac{R^2}{\sigma^2} \left( -\frac{\ell^4}{\sigma^4} d\tau^2 + 2\xi'(\sigma) d\tau d\sigma + d\sigma^2 + d\sigma_2^2 \right) \quad . \quad (4.46)$$

Using this in the free energy (2.12) and solving the resultant equation of motion we find the solution

$$\xi'(\sigma) = \left( \frac{\sigma^8 K(\sigma)^2}{\ell^4 \tilde{\sigma}_0^4 K(\sigma_0)^2} - \frac{\sigma^4}{\ell^4} \right)^{-\frac{1}{2}} \quad , \quad (4.47)$$

where the function  $K(\sigma)$  has been defined in (4.34). Using now the expansion (4.35) we find the length

$$\mathbf{LT}^{-\frac{1}{3}} \left( \frac{3}{2\pi} \right)^{\frac{1}{3}} = -\frac{\sqrt{\pi} \Gamma(\frac{1}{3})}{\Gamma(-\frac{1}{6}) \tilde{\sigma}_0} - \frac{\Gamma(\frac{1}{3}) \Gamma(\frac{7}{6}) \Gamma(\frac{4}{3}) + 2\pi \Gamma(\frac{11}{6})}{18\sqrt{\pi} \Gamma(\frac{4}{3})} \tilde{\sigma}_0^8 \sqrt{\bar{\kappa}} + \mathcal{O}(\bar{\kappa}) \quad . \quad (4.48)$$

As we can see, the dimensionless distance  $\mathbf{LT}^{-\frac{1}{3}}$  is proportional to  $\tilde{\sigma}_0^{-1}$  to leading order, meaning that similarly to the black strings with lightlike separated boundary endpoints analysed in Sec. 3.3, the distance  $\mathbf{LT}^{-\frac{1}{3}}$  decreases for increasing bulk depth  $\tilde{\sigma}_0$ . In fact, numerically we can obtain how the distance  $\mathbf{LT}^{-\frac{1}{3}}$  depends on  $\tilde{\sigma}_0$  for small values of  $\bar{\kappa}$ , which is depicted in the figure below.

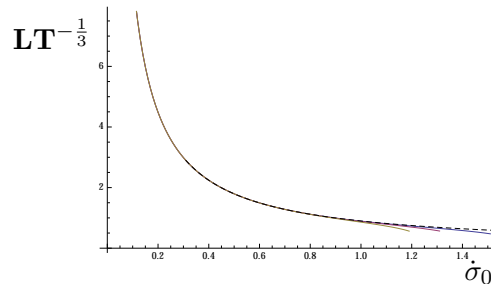


Figure 16: The dimensionless distance  $\mathbf{LT}^{-\frac{1}{3}}$  as a function of  $\tilde{\sigma}_0$  obtained numerically for  $\bar{\kappa} = 0.01$  (yellow curve),  $\bar{\kappa} = 0.001$  (red curve) and  $\bar{\kappa} = 0.0001$  (blue curve). The black dashed line represents  $\bar{\kappa} = 0$ .

As seen from the plot above, the difference between this configuration and those which have been previously analysed, is that the dimensionless distance  $\mathbf{LT}^{-\frac{1}{3}}$  becomes unbounded as  $\tilde{\sigma}_0 \rightarrow 0$ . This means that in order to probe the effect of having a small bulk depth  $\tilde{\sigma}_0$  one must study the behaviour of the free

energy as  $\mathbf{L}\mathbf{T}^{-\frac{1}{3}} \rightarrow \infty$ . We now look at the regularised free energy for these configurations, which can be written as

$$\mathcal{F}_{(3)}[\xi(\tilde{\sigma})] = c_n \tilde{\mathbf{L}} N_{(2)}^{\frac{3}{2}} \sqrt{\lambda_M} \mathbf{T}^{\frac{4}{3}} \left( -\frac{1}{4\tilde{\sigma}_0^4} + \int_0^{\tilde{\sigma}_0} \frac{d\tilde{\sigma}}{\tilde{\sigma}^5} \left( (1-X)\sqrt{1+\tilde{\sigma}^4 \xi'(\tilde{\sigma})^2} - 1 \right) \right) , \quad (4.49)$$

where  $c_n$  has been defined below Eq. (4.37). Expanding this for small  $\bar{\kappa}$  and large  $\mathbf{L}$  we find

$$\mathcal{F}_{(3)} = c_n \tilde{\mathbf{L}} N_{(2)}^{\frac{3}{2}} \sqrt{\lambda_M} \mathbf{L}^4 \left( \frac{19683 \times 3^{\frac{5}{6}} \Gamma(-\frac{1}{6})^2}{20000 \pi^{\frac{4}{3}} \Gamma(-\frac{5}{3})^3 \Gamma(\frac{1}{3})} + C_3 \sqrt{\bar{\kappa}} \mathbf{L}^{-9} \mathbf{T}^3 \right) + \mathcal{O}(\bar{\kappa}) + \mathcal{O}(\mathbf{L}^{-10}) , \quad (4.50)$$

where  $C_3$  is a real constant defined in App. A. The behaviour exhibited here also differs from the previous configurations. In particular, in the case  $\bar{\kappa} = 0$  the free energy is independent of  $T$  and increasing for increasing distance  $\mathbf{L}$ . Going to higher orders in  $\bar{\kappa}$ , for which the details are given in App. A, we can depict the solution space for these solutions. The result is given in the figures below.

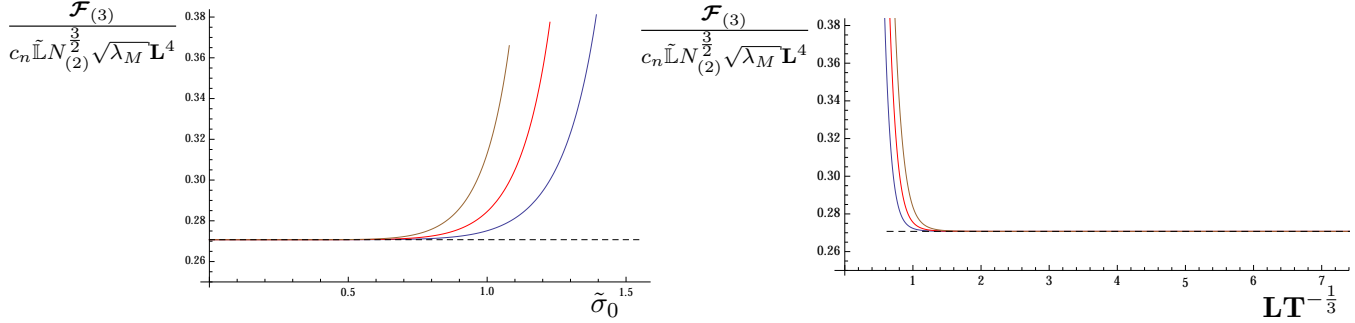


Figure 17: On the left we have plotted the free energy  $\mathcal{F}_{(3)}$  as a function of  $\tilde{\sigma}_0$  and on the right as a function of the distance  $\mathbf{L}\mathbf{T}^{-\frac{1}{3}}$ . The colour coding is the same as in the previous figure.

Noting that the Polyakov surface has zero free energy, then the figure above represents the energy difference between these two configurations. Since this difference is always positive, one concludes that this configuration is not the preferred one with respect to the Polyakov surface. Since now we have several competing configurations at a given temperature  $T$  it is interesting to compare which of the configurations is the preferred one. The free energies (3.36) and (4.49) have been regularised via the same background subtraction procedure so it is sensible to compare the two. Below we show the behaviour of the two free energies for  $\bar{\kappa} = 0.0001$  on the left while on the right we show the behaviour of  $\Delta\mathcal{F}_{(13)} = \mathcal{F}_{(1)} - \mathcal{F}_{(3)}$ .

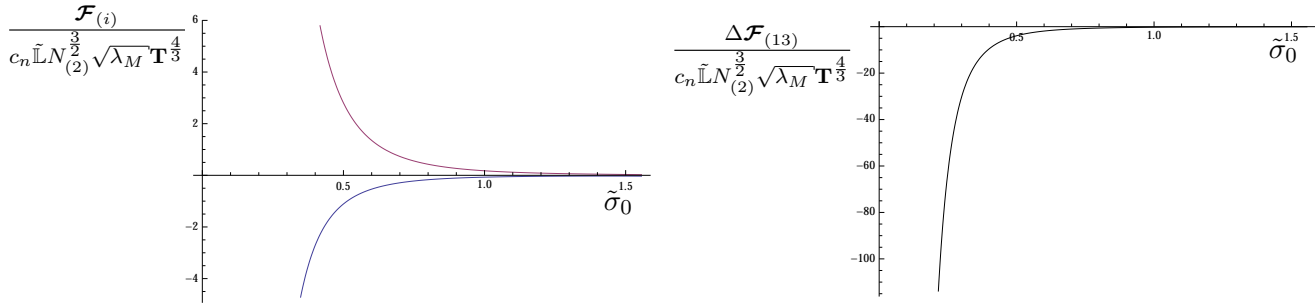


Figure 18: On the left as we display the free energy  $\mathcal{F}_{(1)}$  given by (4.37) (blue curve) and  $\mathcal{F}_{(3)}$  given by (4.49) (red curve) as a function of  $\tilde{\sigma}_0$  for  $\bar{\kappa} = 0.0001$ . On the right we have the free energy difference  $\Delta\mathcal{F}_{(13)}$  as a function of  $\tilde{\sigma}_0$  for  $\bar{\kappa} = 0.0001$ .

We see from the figures above that the free energy (4.37) is always negative while (4.49) is always positive. The free energy difference satisfies  $\Delta\mathcal{F}_{(13)} < 0$  and therefore the configuration with free energy  $\mathcal{F}_{(1)}$  given by (4.37) is always the preferred configuration among the two for a given  $\tilde{\mathbb{L}}, \mathbf{T}, N_{(2)}, \lambda_M, \tilde{\sigma}_0$  and  $\bar{\kappa}$ .

## 5 Discussion

In this paper we have constructed several new configurations of stretched black strings and black membranes in the probe approximation both in  $\text{AdS}_5 \times S^5$ ,  $\text{Sch}_5 \times S^5$  and in M-theory. According to the method we used, namely the blackfold approach, these solutions are valid supergravity solutions for which their metrics can be constructed perturbatively using a matched asymptotic expansion which takes into account back reaction effects. This paper has provided the first application of this method to the case of Schrödinger backgrounds.

In the case of black strings in  $\text{AdS}_5 \times S^5$ , we have reviewed previously constructed configurations in a black hole background [26] and generalised it to thermal AdS. These configurations represent Wilson loop operators at finite temperature in the dual gauge theory at strong 't Hooft coupling. In this context the free energy of the black string yields the quark-antiquark potential which scales with the number of fundamental strings  $k$  and is inversely proportional to the length  $L$  of the pair. Following the black probe method leads to dominant finite temperature corrections in the quark-antiquark potential and to new physics, which can be observed already in thermal AdS, without the presence of a black hole.

In Sec. 3 we have constructed black strings in  $\text{Sch}_5 \times S^5$ , whose near horizon geometry is composed of black, asymptotically flat, fundamental strings. In this setting we constructed the analogous configuration of the one found in  $\text{AdS}_5 \times S^5$  and compared the two cases. We found that there exists a string configuration that interpolates between the AdS Wilson loop and a new configuration in  $\text{Sch}_5 \times S^5$ . Due to the presence of the null direction in  $\text{Sch}_5 \times S^5$  we also saw that it was possible, not only to have a configuration analogous to the one found in  $\text{AdS}_5 \times S^5$ , but also to have a black string configuration in which its boundary endpoints are lightlike separated. This configuration turns out not to be the preferred one, having a higher value of its free energy. The differences between null AdS spaces and Schrödinger spacetimes are subtle, namely when probing the spacetimes with particle or field probes [12], but here we have seen that the black probe method allows to distinguish between these two spacetimes by placing probes along the null direction of  $\text{Sch}_5 \times S^5$  and hence leading to a configuration with no AdS counterpart.

According to the works of [3, 4, 5], backgrounds obtained via the null Melvin twist of supergravity branes should have well defined quantum field theory duals. In particular, in the case of  $\text{Sch}_5 \times S^5$  studied here, the dual quantum field theory is supposed to be obtained via the  $\text{DLCQ}_\beta$  of  $\mathcal{N} = 4$  SYM and to lead to a non-relativistic CFT. In this line of thought it is relevant to ask if the black string geometries that we constructed here may be considered to be dual to Wilson loop operators defined in the gauge theory. In this context we note that the interpolating black string geometry in  $\text{Sch}_5 \times S^5$  in Sec. 3.1 analogous to the one in  $\text{AdS}_5 \times S^5$ , has a free energy which scales with the total number of fundamental strings  $k$  but now is inversely proportional to a certain power of  $L$  of the quark-antiquark pair which can vary between  $L^{-1}$  to  $L^{-2}$  depending on the magnitude of the deformation parameter  $\ell$  and for large enough  $\ell$  it becomes proportional to  $\ell$  itself. We have seen that for the case of the configurations of Sec. 3.2, which connect at large  $\ell$  to the interpolating strings, the combination  $\mathcal{F}_{\text{loop}} L^2 / \ell$  appearing in the free



energy (3.36) is scale invariant and only depends on the scale invariant combination  $L(T/\ell)^{1/2}$ . This indeed suggests that the dual gauge theory is conformally invariant as previously suspected [3, 4, 5] and furthermore that the deformation parameter  $\ell$  has a physical role to play in the gauge theory. It is worth noting that the blackness of the probe, characterised by corrections in powers of  $\kappa$ , enters at the same order in the quark-antiquark potential as in its AdS counterpart and it is as sensitive to temperature corrections, which enter at order  $T^3$ , but more sensitive to small distances between the quark and the antiquark, scaling with  $L^{-2}$  at zero temperature. One may also wonder about the dual interpretation of the black strings in  $\text{Sch}_5 \times S^5$  with boundary endpoints lightlike separated. However, the issue of whether or not the null direction should be compactified for the purposes of holography is not yet settled, which leads to lack of intuition about the role of these configurations. Related to this, the exact details of the dual field theory action obtained by the DLCQ $_\beta$  of  $\mathcal{N} = 4$  SYM are still lacking but they would consequently allow for the study of expectation values of Wilson loop operators at zero temperature and to check if the corresponding result can be matched with the predictions made here. This endeavour would lead to a better understanding of holography in these spacetimes.

In Sec. 4 we have studied the case of stretched black membranes in  $\text{AdS}_7 \times S^4$  and found that, similarly to the case of  $\text{AdS}_5 \times S^5$ , the black probe method leads to new physics and dominant contributions to the free energy which were not seen in the work of [32]. In particular, we see that there is a contribution of order  $(TL)^3$  in the free energy of the black membrane (4.15) which is dominant over the contribution due to the black hole background entering at order  $(TL)^6$ . Since in this background the black membrane is less sensitive to the corrections due to the black hole, corrections due to the blackness of the probe become even more relevant and, for example, change the qualitatively behaviour of the onset of the Debye screening effect. In  $\text{Sch}_7 \times S^4$  we constructed analogous configurations and also three other black membrane configurations, depending on how we place its endpoints on the boundary geometry, and which have no AdS counterpart. We have in particular seen that the interpolating black membrane with an AdS counterpart, the free energy scales with the number of probe N2-branes  $N_{(2)}^{\frac{3}{2}}$  and with a power of  $L$  that varies from  $L^{-2}$  to  $L^{-4}$  as the deformation parameter  $\ell$  is increased continuously to higher and higher values. Again, assuming the existence of a dual gauge theory and that these configurations represent Wilson surfaces, we see that the free energy (4.38) suggests that the dual non-relativistic gauge theory is conformally invariant. One may also wonder if it is possible to check this dependence in the expectation values of Wilson surface operators in  $\text{Sch}_7 \times S^4$  as it was recently done in [42] for the same operators in  $\text{AdS}_7 \times S^4$ .

We also note that all the black strings and black membranes constructed in Schrödinger spacetimes in this paper were done in thermal Schrödinger, without the presence of a black hole. Schrödinger black branes can be obtained via the null Melvin twist of supergravity black branes but generate non-trivial dilatonic fields [3, 4, 5]. Within the blackfold approach, coupling black probes to dilatonic background fields has not been yet properly understood, though currently under study [38]. Therefore, we have not considered such background geometries in this work but we admit that it would be worthwhile studying them as one would most likely observe non-trivial phase transitions and Debye screening effects. It is worth mentioning that all the configurations that we constructed both in thermal AdS and thermal Schrödinger do not exhibit Debye screening effects, suggesting that it would be interesting to understand whether a general argument ruling out the existence of these effects exists.

Finally, we note that applying the null Melvin twist on supergravity branes leads to black branes whose

near horizon geometry exhibits the Schrödinger symmetry group [3, 4, 5, 16]. It would be interesting to take these branes and use them to probe the spacetime and in the process construct new black hole geometries with non-relativistic near-horizon geometry. We leave this possibility for future studies.

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## A Higher order expansions for black membranes in Schrödinger

In this appendix we collect some details on the solutions presented in Sec. 4.2. In particular we give higher order expansions in  $\bar{\kappa}$  for the dimensionless distances and free energies characterising each solution and which allowed us to analyse the solution space.

### Membranes with boundary strings spatially extended and spacelike separated

We have solved for the dimensionless distance numerically and seen that it is well approximated for small  $\bar{\kappa}$  and up to values  $\sigma \sim \sigma_c$  by the expansion.

$$\begin{aligned} \mathbf{LT}^{\frac{1}{3}} \left( \frac{\pi}{12} \right)^{1/3} &= \frac{\sqrt{\pi} \Gamma(\frac{3}{5})}{\Gamma(\frac{1}{10})} \tilde{\sigma}_0 + \frac{\pi}{30} \left( \frac{2\Gamma(\frac{3}{5})}{\sqrt{\pi} \Gamma(\frac{1}{10})} - 1 \right) \sqrt{\bar{\kappa}} \tilde{\sigma}_0^{10} + \frac{\pi \left( \frac{\frac{276\Gamma(\frac{8}{5})}{\Gamma(\frac{1}{10})} + \frac{11\Gamma(\frac{17}{5})}{\Gamma(\frac{19}{10})} - 96 \right)}{4320} \bar{\kappa} \tilde{\sigma}_0^{19} \\ &+ \frac{\sqrt{\pi} \left( -346500\sqrt{\pi} + \frac{950675\Gamma(\frac{8}{5})}{\Gamma(\frac{1}{10})} + \frac{193116\Gamma(\frac{7}{5})}{\Gamma(\frac{19}{10})} - \frac{4650\Gamma(\frac{43}{10})}{\Gamma(\frac{14}{5})} \right)}{2 \times 8910000} \bar{\kappa}^{3/2} \tilde{\sigma}_0^{28} + \mathcal{O}(\bar{\kappa}^2) . \end{aligned} \quad (\text{A.1})$$

Using the above expansion and a similar one for the free energy (4.37),

$$\begin{aligned} \frac{\mathcal{F}_{(1)}}{c_n \tilde{\mathbf{L}} N_{(2)}^{\frac{3}{2}} \sqrt{\lambda_M} \mathbf{T}^{\frac{4}{3}}} &= \frac{\sqrt{\pi} \Gamma(-\frac{2}{5})}{10\Gamma(\frac{1}{10}) \tilde{\sigma}_0^4} - \frac{1}{135} \sqrt{\pi} \left( 9\sqrt{\pi} + \frac{10\Gamma(\frac{3}{5})}{\Gamma(-\frac{9}{10})} \right) \sqrt{\bar{\kappa}} \tilde{\sigma}_0^5 \\ &- \frac{\sqrt{\pi} \left( 720\sqrt{\pi} - \frac{418\Gamma(\frac{7}{5})}{\Gamma(\frac{9}{10})} - \frac{147\Gamma(\frac{8}{5})}{\Gamma(\frac{11}{10})} \right)}{32400} \bar{\kappa} \tilde{\sigma}_0^{14} \\ &- \frac{\sqrt{\pi} \left( 61200\sqrt{\pi} + \frac{8463\Gamma(\frac{13}{10})}{\Gamma(\frac{4}{5})} - \frac{46816\Gamma(\frac{7}{5})}{\Gamma(\frac{9}{10})} - \frac{13422\Gamma(\frac{8}{5})}{\Gamma(\frac{11}{10})} \right)}{3888000} \bar{\kappa}^{3/2} \tilde{\sigma}_0^{23} + \mathcal{O}(\bar{\kappa}^2) , \end{aligned} \quad (\text{A.2})$$

where  $c_n$  was defined below (4.37), we are able to analyse the solution space. Inverting (A.1) and introducing in (A.2) leads to the value of  $C_1$  introduced in (4.38),

$$C_1 = \frac{25 \left( 9\sqrt{\pi}\Gamma\left(-\frac{9}{10}\right) + 10\Gamma\left(\frac{3}{5}\right) \right) \Gamma\left(\frac{11}{10}\right)^4}{9 \times 3^{\frac{2}{3}}(2\pi)^{\frac{1}{3}}\Gamma\left(\frac{3}{5}\right)^5} . \quad (\text{A.3})$$

### Membranes with boundary strings null extended and spacelike separated

For these configurations we have also studied numerically the dimensionless distance and seen that it is well approximated by the following expansion,

$$\begin{aligned} \mathbf{L}\mathbf{T}^{\frac{1}{3}}\left(\frac{\pi}{12}\right)^{1/3} &= \frac{\sqrt{\pi}\Gamma\left(\frac{2}{3}\right)}{\Gamma\left(\frac{1}{6}\right)}\tilde{\sigma}_0 + \frac{\Gamma\left(\frac{2}{3}\right)\Gamma\left(\frac{5}{6}\right) + \sqrt{3}\Gamma\left(-\frac{2}{3}\right)\Gamma\left(\frac{13}{6}\right)}{18\sqrt{\pi}}\sqrt{\bar{\kappa}}\tilde{\sigma}_0^{10} \\ &+ \sqrt{\pi}\frac{\frac{82779\Gamma\left(\frac{2}{3}\right)}{\Gamma\left(\frac{1}{6}\right)} - \frac{25480\Gamma\left(\frac{7}{6}\right)}{\Gamma\left(\frac{5}{3}\right)}}{2 \times 88452}\bar{\kappa}\tilde{\sigma}_0^{19} + \sqrt{\pi}\frac{\frac{24316558464\Gamma\left(\frac{2}{3}\right)}{\Gamma\left(\frac{1}{6}\right)} - \frac{6522475505\Gamma\left(\frac{7}{6}\right)}{\Gamma\left(\frac{5}{3}\right)}}{2 \times 13450364928}\bar{\kappa}^{3/2}\tilde{\sigma}_0^{28} + \mathcal{O}(\bar{\kappa}^2) . \end{aligned} \quad (\text{A.4})$$

Expanding the free energy (4.43) in powers of  $\bar{\kappa}$ ,

$$\begin{aligned} \frac{\mathcal{F}_{(2)}}{c_s \tilde{\mathbb{L}} N^{\frac{3}{2}}_{(2)} \sqrt{\lambda_M} \mathbf{T}^{\frac{2}{3}}} &= \frac{\sqrt{\pi}\Gamma\left(-\frac{1}{3}\right)}{6\Gamma\left(\frac{1}{6}\right)\tilde{\sigma}_0^2} + \frac{\sqrt{\pi}\left(\frac{2\Gamma\left(\frac{2}{3}\right)^2}{\Gamma\left(\frac{1}{6}\right)} - 5\Gamma\left(\frac{7}{6}\right)\right)}{18\Gamma\left(\frac{2}{3}\right)}\sqrt{\bar{\kappa}}\tilde{\sigma}_0^7 \\ &+ \frac{\sqrt{\pi}\left(\frac{92823\Gamma\left(\frac{2}{3}\right)}{\Gamma\left(\frac{1}{6}\right)} - \frac{25480\Gamma\left(\frac{7}{6}\right)}{\Gamma\left(\frac{5}{3}\right)}\right)}{176904}\bar{\kappa}\tilde{\sigma}_0^{16} + \frac{\sqrt{\pi}\left(\frac{29744167200\Gamma\left(\frac{2}{3}\right)}{\Gamma\left(\frac{1}{6}\right)} - \frac{7686566251\Gamma\left(\frac{7}{6}\right)}{\Gamma\left(\frac{5}{3}\right)}\right)}{33625912320}\bar{\kappa}^{3/2}\tilde{\sigma}_0^{25} + \mathcal{O}(\bar{\kappa}^2) , \end{aligned} \quad (\text{A.5})$$

where  $c_s$  was defined below (4.43), allows us to probe the solution space. Inverting (A.4) and using it in (A.5) allows us to derive the constant  $C_2$  introduced in (4.44),

$$C_2 = \frac{3^{\frac{1}{6}}\Gamma\left(\frac{1}{6}\right)^7\left(45 \times 2^{\frac{1}{3}}\sqrt{3\pi}\Gamma\left(\frac{7}{6}\right) - 7\Gamma\left(-\frac{2}{3}\right)^2\right)}{64\pi^{7/6}\Gamma\left(-\frac{1}{3}\right)^7\Gamma\left(\frac{2}{3}\right)} . \quad (\text{A.6})$$

### Membranes with boundary strings spatially extended and lightlike separated

At the end of Sec. 4.2 we have studied numerically the dimensionless length  $\mathbf{L}$ . We have checked that its behaviour is well approximated for small  $\bar{\kappa}$  by the expansion

$$\begin{aligned} \mathbf{L}\mathbf{T}^{-\frac{1}{3}}\left(\frac{3}{2\pi}\right)^{\frac{1}{3}} &= -\frac{\sqrt{\pi}\Gamma\left(\frac{1}{3}\right)}{\Gamma\left(-\frac{1}{6}\right)\tilde{\sigma}_0} - \frac{\left(\Gamma\left(\frac{1}{3}\right)\Gamma\left(\frac{7}{6}\right)\Gamma\left(\frac{4}{3}\right) + 2\pi\Gamma\left(\frac{11}{6}\right)\right)}{18\sqrt{\pi}\Gamma\left(\frac{4}{3}\right)}\sqrt{\bar{\kappa}}\tilde{\sigma}_0^8 \\ &+ \frac{\sqrt{\pi}\left(4400\Gamma\left(-\frac{1}{6}\right)^2 - 13713\Gamma\left(\frac{1}{3}\right)^2\right)}{2 \times 53460\Gamma\left(-\frac{1}{6}\right)\Gamma\left(\frac{1}{3}\right)}\bar{\kappa}\tilde{\sigma}_0^{17} \\ &+ \left(\frac{56722777231\Gamma\left(\frac{7}{6}\right)\Gamma\left(\frac{4}{3}\right)}{2\sqrt{\pi} \times 58193562240} - \frac{151202389099\sqrt{\frac{1}{3}}\Gamma\left(\frac{5}{6}\right)\Gamma\left(\frac{5}{3}\right)}{2\sqrt{\pi} \times 92748644352}\right)\bar{\kappa}^{3/2}\tilde{\sigma}_0^{26} + \mathcal{O}(\bar{\kappa}^2) . \end{aligned} \quad (\text{A.7})$$

A similar expansion in the free energy (4.49),

$$\begin{aligned} \frac{\mathcal{F}_{(3)}}{c_n \tilde{L} N_{(2)}^{\frac{3}{2}} \sqrt{\lambda_M} \mathbf{T}_{\frac{4}{3}}} &= \frac{\sqrt{\pi} \Gamma(-\frac{2}{3})}{6 \Gamma(-\frac{1}{6}) \tilde{\sigma}_0^4} - \frac{\left(8\pi \Gamma(\frac{5}{6}) + \Gamma(\frac{1}{3})^2 \Gamma(\frac{7}{6})\right)}{18 \sqrt{\pi} \Gamma(\frac{1}{3})} \sqrt{\bar{\kappa}} \tilde{\sigma}_0^5 \\ &- \frac{\sqrt{\pi} \left(\frac{19869 \Gamma(\frac{1}{3})}{\Gamma(-\frac{1}{6})} + \frac{8800 \Gamma(\frac{5}{6})}{\Gamma(\frac{4}{3})}\right)}{106920} \bar{\kappa} \tilde{\sigma}_0^{14} - \frac{\sqrt{\pi} \left(\frac{1645699776 \Gamma(\frac{1}{3})}{\Gamma(-\frac{1}{6})} + \frac{629014375 \Gamma(\frac{5}{6})}{\Gamma(\frac{4}{3})}\right)}{5173217280} \bar{\kappa}^{3/2} \tilde{\sigma}_0^{23} + \mathcal{O}(\bar{\kappa}^2) \quad , \end{aligned} \quad (\text{A.8})$$

has allowed us to depict the solution space in Sec. 4.2. Inverting (A.7) and plugging it into (A.8) yields the constant  $C_3$  introduced in (4.50),

$$C_3 = \frac{1309000000 \left(\frac{2}{3}\right)^{2/3} \pi^{11/3} \Gamma(-\frac{20}{3}) \Gamma(-\frac{5}{3})^3 \Gamma(-\frac{5}{6})}{387420489 \Gamma(-\frac{1}{6})^3} \quad . \quad (\text{A.9})$$

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